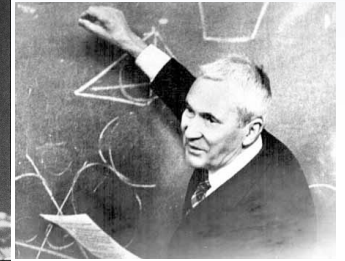
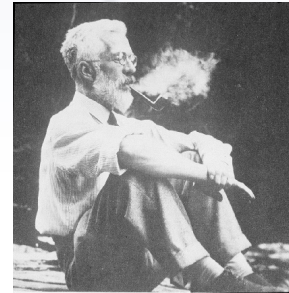
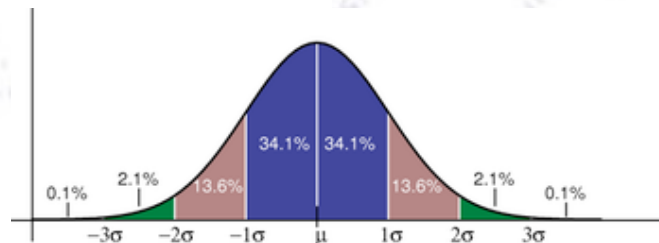


Applied Statistics

Mean and Width



Troels C. Petersen (NBI)

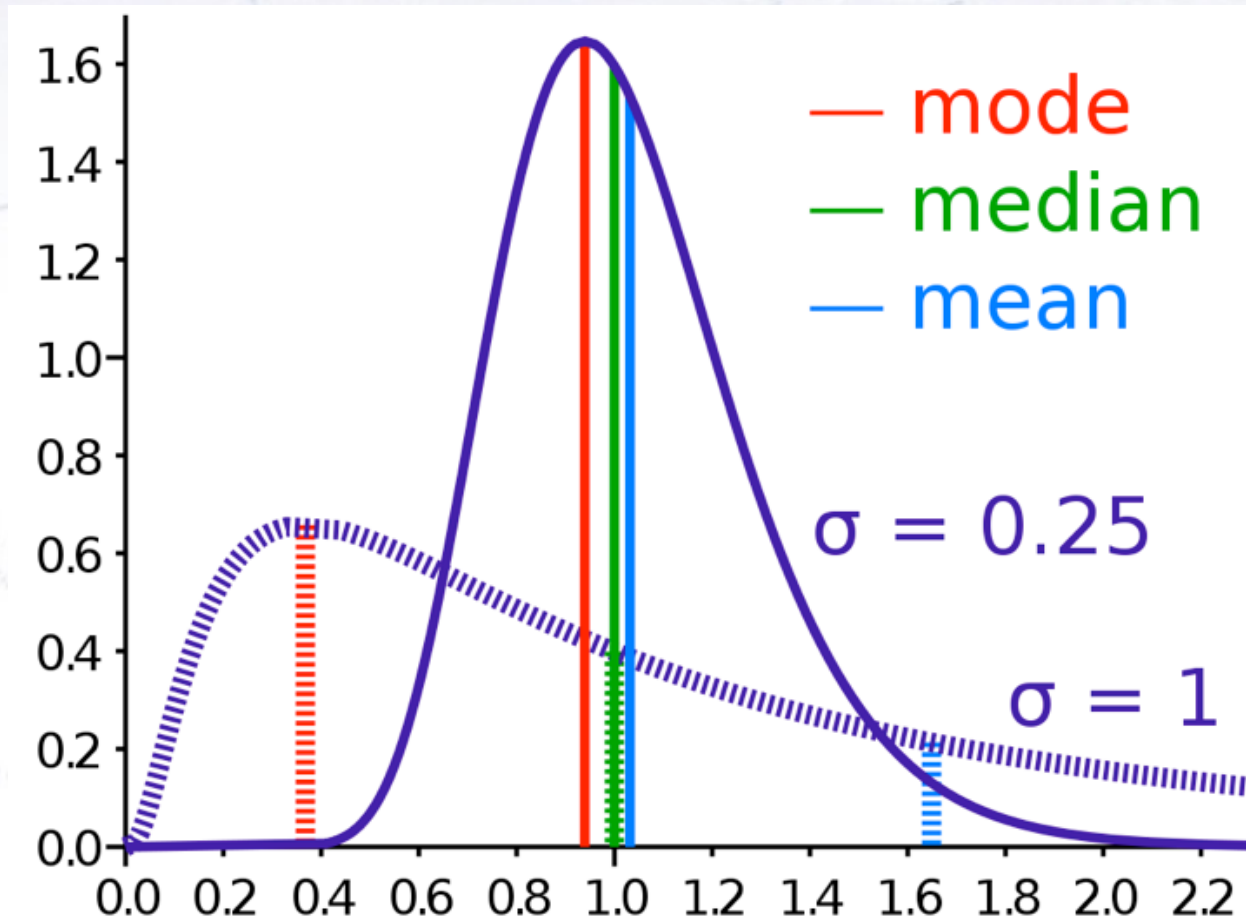


"Statistics is merely a quantisation of common sense"

Defining the mean

There are several ways of defining “a typical” value from a dataset:

- a) Arithmetic mean b) Mode (most probably) c) Median (half below, half above)
d) Geometric mean e) Harmonic mean f) Truncated mean (robustness)



Mean and Width

It turns out, that the best estimator for the **mean** is (as you all know):

$$\hat{\mu} = \frac{1}{N} \sum_i x_i = \bar{x}$$

For the **width** of the distribution (a.k.a. **standard deviation** or **RMS**) it is:

$$\hat{\sigma} = \sqrt{\frac{1}{N} \sum_i (x_i - \mu)^2}$$

Note the “hat”, which means “estimator”. It is sometimes dropped...

Mean and Width

It turns out, that the best estimator for the **mean** is (as you all know):

$$\hat{\mu} = \frac{1}{N} \sum_i x_i = \bar{x}$$

For the **width** of the distribution (a.k.a. **standard deviation** or **RMS**) it is:

$$\hat{s} = \sqrt{\frac{1}{N-1} \sum_i (x_i - \bar{x})^2}$$

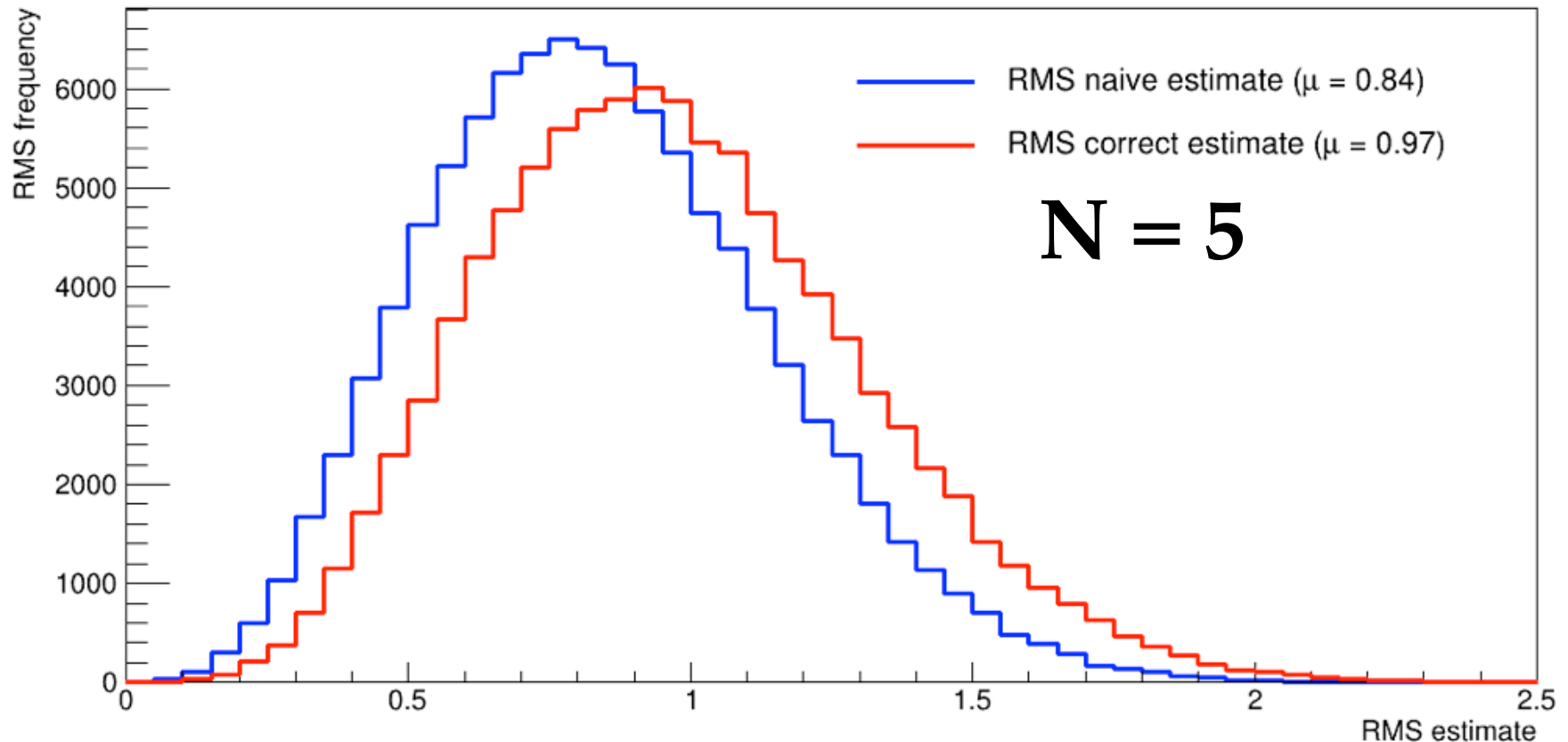
Note the “hat”, which means “estimator”. It is sometimes dropped...

How incorrect is the naive RMS?

Such questions can most easily be answered by a small simulation...

Produce $N=5$ numbers from a unit Gaussian, and calculate the RMS estimate:

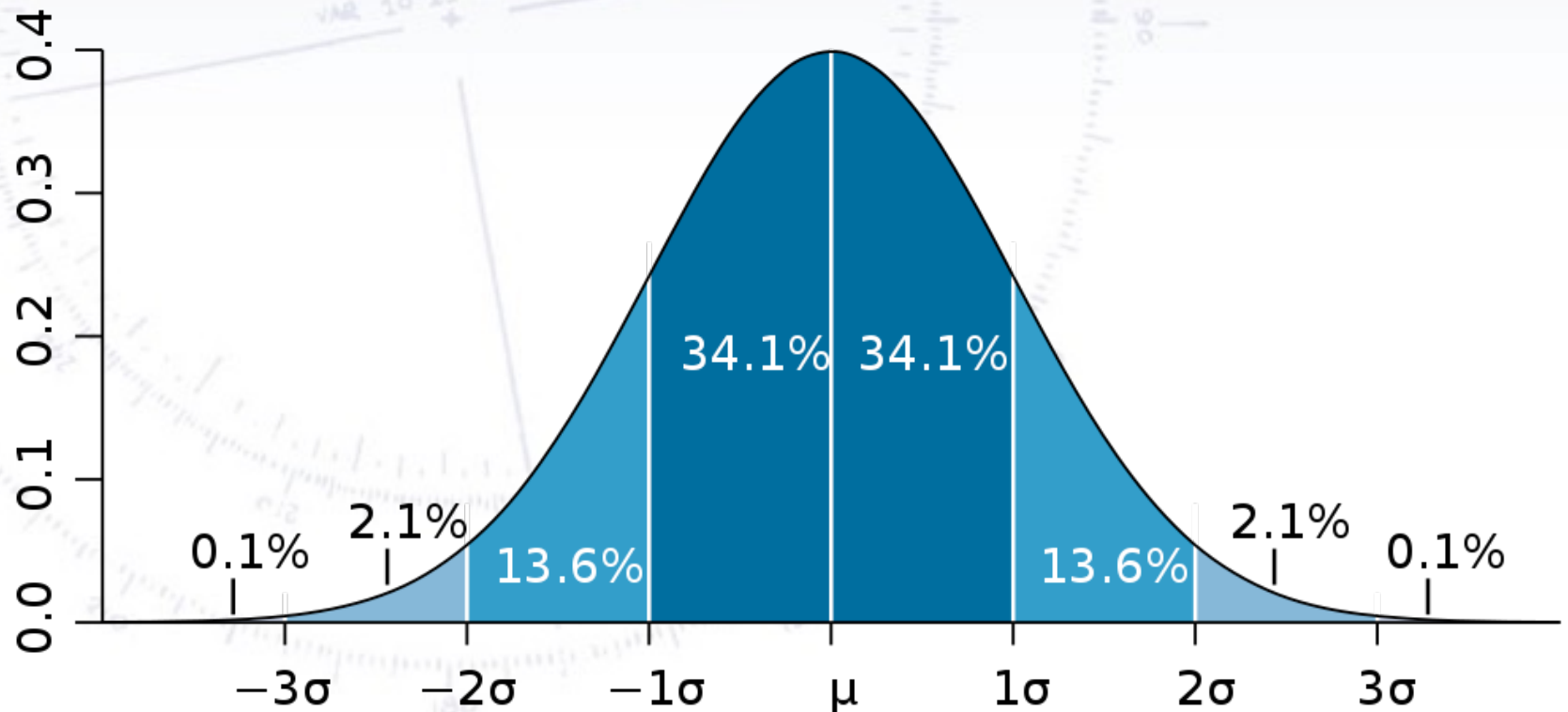
Distribution of RMS estimates on five unit Gaussian numbers



So, the “naive” RMS underestimates the uncertainty a bit...

Relation between RMS and Gaussian width...

When a distribution is Gaussian, the RMS corresponds to the Gaussian width σ :



Mean and Width

What is the **uncertainty on the mean**? And how quickly does it improve with more data?

$$\hat{\sigma}_{\mu} = \hat{\sigma} / \sqrt{N}$$

Example:

Cavendish Experiment

(measurement of Earth's density)

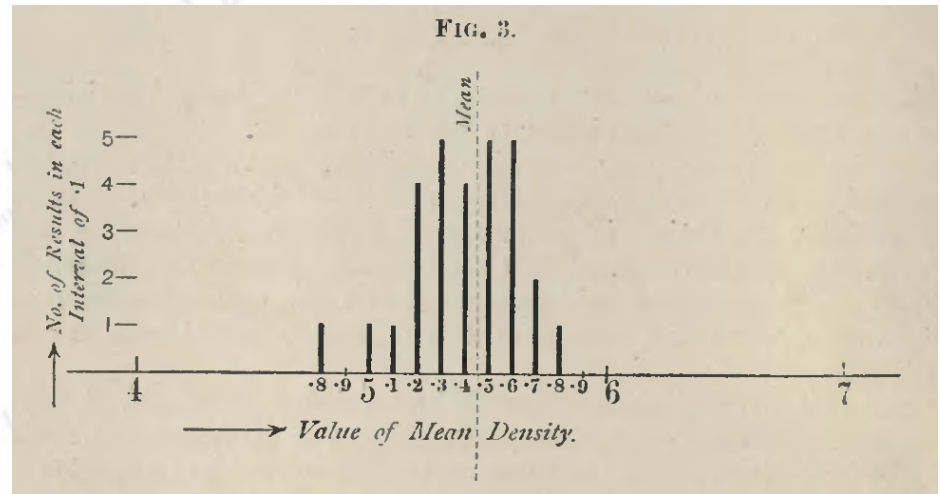
$N = 29$

$\mu = 5.42$

$\sigma = 0.333$

$\sigma(\mu) = 0.06$

Earth density = 5.42 ± 0.06



Mean and Width

What is the **uncertainty on the mean**? And how quickly does it improve with more data?

$$\hat{\sigma}_{\mu} = \hat{\sigma} / \sqrt{N}$$

Please, commit to memory now!

Example.

Cavendish Experiment
(measurement of Earth's density)

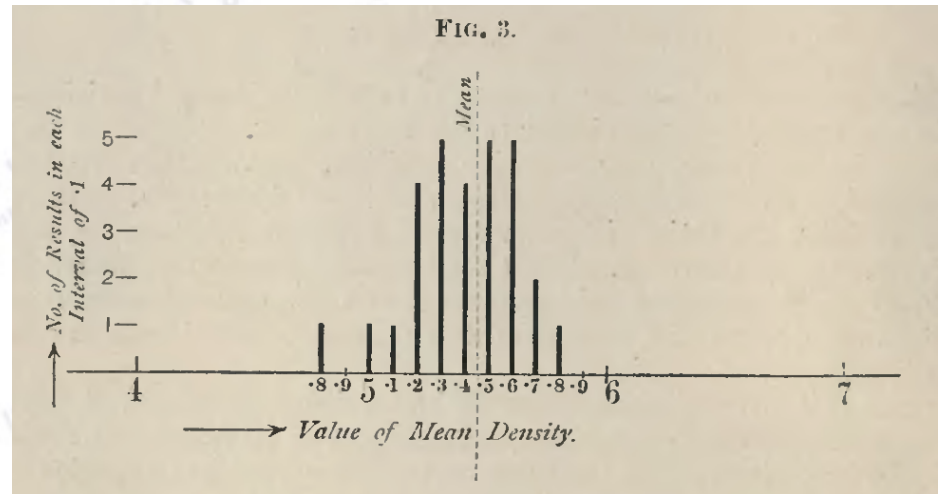
$N = 29$

$\mu = 5.42$

$\sigma = 0.333$

$\sigma(\mu) = 0.06$

Earth density = 5.42 ± 0.06



Weighted Mean

What if we are given data, which has different uncertainties?

How to average these, and what is the uncertainty on the average?

$$\hat{\mu} = \frac{\sum x_i / \sigma_i^2}{\sum 1 / \sigma_i^2}$$

For measurements with varying uncertainty, there is no meaningful RMS!

The uncertainty on the mean is:

$$\hat{\sigma}_{\mu} = \sqrt{\frac{1}{\sum 1 / \sigma_i^2}}$$

Can be understood intuitively, if two persons combine 1 vs. 4 measurements

Resolution using InterQuantile Range

A useful measure of resolution is the InterQuantile Range (IQR), as this is not affected by long tails.

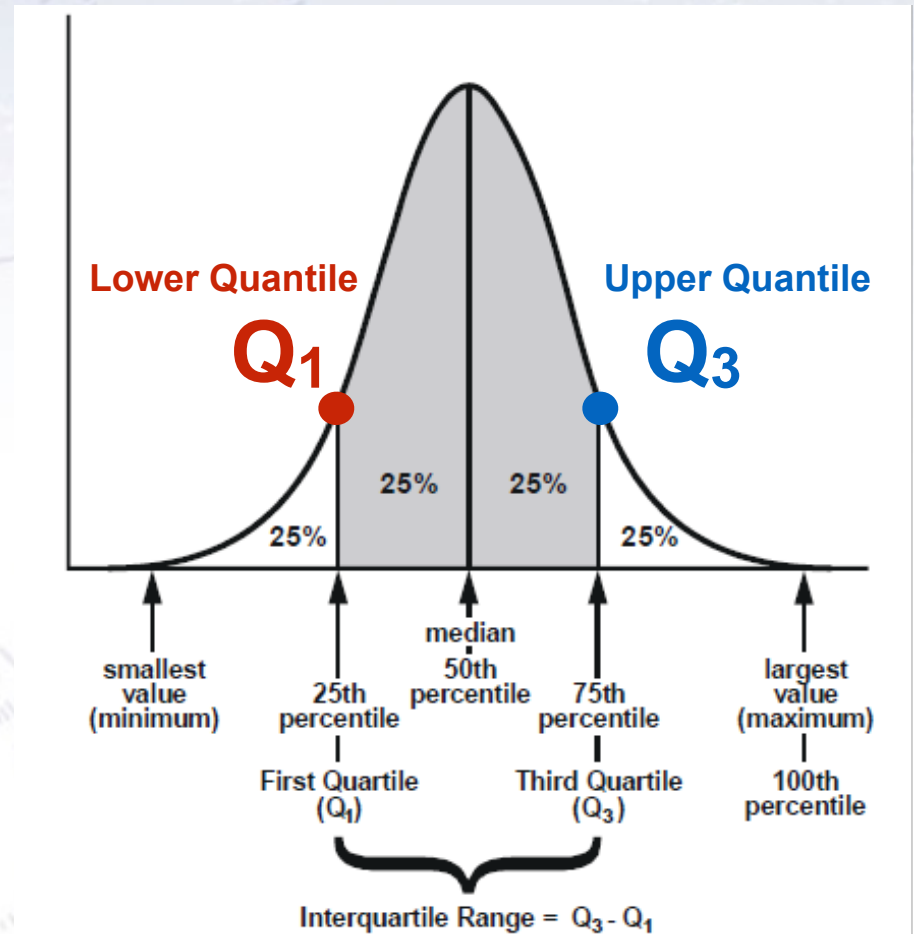
IQR measures **statistical dispersion**, calculated as the difference

$$\text{IQR} = Q_3 - Q_1$$

The InterQuantile Efficiency (IQE) is defined as:

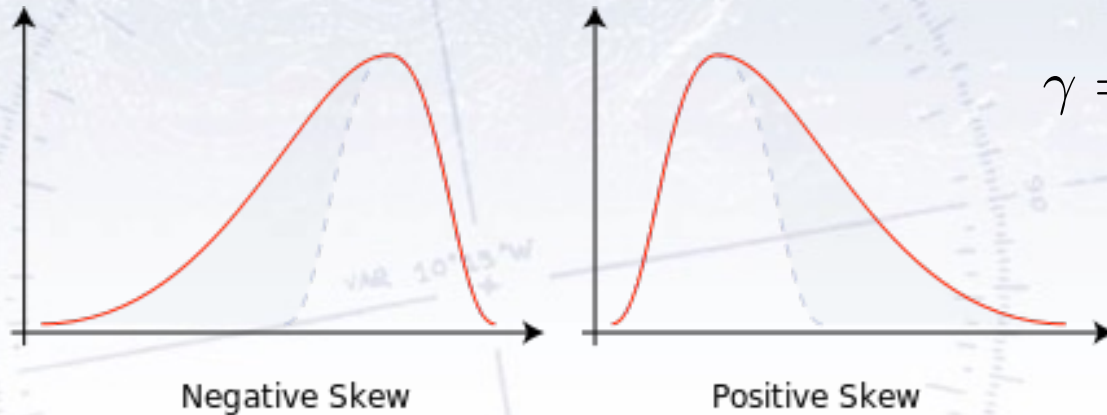
$$\text{IQE} = \text{IQR} / 1.349$$

The factor $1.349 = 2 \Phi^{-1}(0.75)$ ensures that $\text{IQR} = 1$ for a unit Gaussian.



Skewness and Kurtosis

Higher moments reveal something about a distributions asymmetry and tails:



$$\gamma = \frac{\frac{1}{N} \sum_i (x_i - \bar{x})^3}{\left(\frac{1}{N} \sum_i (x_i - \bar{x})^2\right)^{3/2}}$$

$$\kappa = \frac{\frac{1}{N} \sum_i (x_i - \bar{x})^4}{\left(\frac{1}{N} \sum_i (x_i - \bar{x})^2\right)^2} - 3$$

LEPTOKURTIC
(thicker tails)

MESOKURTIC
(normal tails)

PLATYKURTIC
(thinner tails)

