## Applied Statistics <br> Error propagation



Troels C. Petersen (NBI)

"Statistics is merely a quantisation of common sense"

## Error propagation

Applied Statistics
Overview of subjects


## Error propagation

Imagine that $y$ is a function of $x_{i}$

$$
y\left(x_{i}\right)
$$

and that we wish to find the error on $y$ from the errors on $x_{i}$.

$$
\begin{aligned}
& \sigma\left(x_{i}\right)=0.8 \\
& \sigma\left(y\left(x_{i}\right)\right)=?
\end{aligned}
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$$
\begin{array}{r}
\sigma\left(x_{i}\right)=0.8 \\
\sigma\left(y\left(x_{i}\right)\right)= \\
\frac{\partial y}{\partial x_{i}} \times 0.8
\end{array}
$$



## Error propagation

Note, the approximation here:
The derivative of $y-d y / d x_{i}$ - should be relatively constant. If not, the error propagation formula breaks down.
and that we wish to find the error on $y$ from the errors on $x_{i}$.

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\end{gathered}
$$



## Error propagation

Imagine that $y$ is a function of $x_{i}$, and that we wish to find the error on $y$ from the errors on $\mathrm{x}_{\mathrm{i}}$. Making a Taylor expansion of the function y gives:

$$
y(\bar{x}) \simeq y(\bar{\mu})+\sum_{i}^{n} \frac{\partial y}{\partial x_{i}}\left(x_{i}-\mu_{i}\right)
$$

In order to get the uncertainty of $y$ as a function of the variables $x_{i}$ we calculate:

$$
\begin{gathered}
\sigma_{x}^{2}=\overline{x^{2}}-\bar{x}^{2}=E\left[x^{2}\right]-E^{2}[x] \\
E[y(\bar{x})] \\
\simeq y(\bar{\mu}) \\
E\left[y^{2}(\bar{x})\right]
\end{gathered}
$$

## Error propagation formula

Subtracting the two formulae, we obtain:

$$
\sigma_{y}^{2}=\sum^{n}\left[\frac{\partial y}{\partial r_{i}} \frac{\partial y}{\partial r_{i}}\right] \quad V_{i j}
$$

If there are no correlations, only the diagonal (individual errors) enter:

$$
\sigma_{y}^{2}=\sum_{i}^{n}\left[\frac{\partial y}{\partial x_{i}}\right]_{\bar{x}=\bar{y}}^{2} \sigma_{i}^{2}
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## Error propagation formula

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$$

## Specific error propagation formula Addition

Specific formula:

$$
x=u+v
$$

$$
\sigma_{x}^{2}=\sigma_{u}^{2}+\sigma_{v}^{2}+2 V_{u v}
$$

General formula:

$$
\begin{gathered}
x=a u+b v \\
\sigma_{x}^{2}=a^{2} \sigma_{u}^{2}+b^{2} \sigma_{v}^{2}+2 a b V_{u v}
\end{gathered}
$$

"When adding numbers, their errors add in quadrature"

## Specific error propagation formula Multiplication

$$
\begin{gathered}
x=u v \\
\sigma_{x}^{2}=\left(v \sigma_{u}\right)^{2}+\left(u \sigma_{v}\right)^{2}+2 u v V_{u v}
\end{gathered}
$$

Dividing by $\mathrm{x}^{2}$ to get relative terms, we obtain:

$$
\frac{\sigma_{x}^{2}}{x^{2}}=\frac{\sigma_{u}^{2}}{u^{2}}+\frac{\sigma_{v}^{2}}{v^{2}}+2 \frac{V_{u v}}{u v}
$$

"When multiplying numbers, their RELATIVE errors add in quadrature"

## Error propagation at work...



John Harrison (24 March 1693-24 March 1776)
British clockmaker extraordinaire "Won" the Longitude Act prize (3 sec/day).

Harrison's first sea clock (H1)


Harrison build H1-H5. K1 (Copy of H4) was used by James Cook.

## Error propagation at work...

Harrison's Gridiron pendulum is designed to cancel the change in length (in fact moment of inertia) with temperature.


Coefficient of thermal expansion: Iron $=11.8 \times 10^{-6} / \mathrm{C}^{\circ} \mathrm{Zinc}=30.2 \times 10^{-6} / \mathrm{C}^{\circ}$

## Error propagation at more work．．．

Analysis of tiny differences in Uranus＇orbit from Newtonian prediction led to the prediction and discovery of Neptune！

## Continuing with Mercury．．．

Table II．Contributions to the motion of the perihelia of Mercury and the earth．

| Cause |  | Motion of perihelion |  |
| :---: | :---: | :---: | :---: |
|  | $m^{-1}$ | Mercury | Earth |
| Mercury | $6000000 \pm 1000000$ | $0.1025 \pm 0$＂． 00 | $-13^{\prime \prime} .75 \pm 2{ }^{\prime \prime} .3$ |
| Venus | $408000 \pm 1000$ | $277.856 \pm 0.68$ | $345.49 \pm 0.8$ |
| Earth | 329390 士 300 | $90.038 \pm 0.08$ |  |
| Mars | $3088000 \pm 3000$ | $2.536 \pm 0.00$ | $97.69 \pm 0.1$ |
| Jupiter | $1047.39 \pm \quad 0.03$ | $153.584 \pm 0.00$ | $696.85 \pm 0.0$ |
| Saturn | 3499 士 4 | $7.302 \pm 0.01$ | $18.74 \pm 0.0$ |
| Uranus | 22800 土 300 | $0.141 \pm 0.00$ | $0.57 \pm 0.0$ |
| Neptune | $19500 \pm 300$ | $0.042 \pm 0.00$ | $0.18 \pm 0.0$ |
| Solar oblateness |  | $0.010 \pm 0.02$ | $0.00 \pm 0.0$ |
| Moon |  |  | $7.68 \pm 0.0$ |
| General precession（Julian century，1850） |  | $5025.645 \pm 0.50$ | $5025.65 \pm 0.5$ |
| Sum <br> Observed motion |  | $5557.18 \pm 0.85$ | $6179.1 \pm 2.5$ |
|  |  | $5599.74 \pm 0.41$ | $6183.7 \pm 1.1$ |
| Difference <br> Relativity effect |  | $42.56 \pm 0.94$ | $4.6 \pm 2.7$ |
|  |  | $43.03 \pm 0.03$ | $3.8 \pm 0.0$ |



## Simulating error propagation

Imagine that y is a very complicated function of $\mathrm{x}_{\mathrm{i}}$, perhaps not even parametric (i.e. not a function, but rather a model).

A simple method is to use simulation:

- Choose random values of $x_{i}$, corresponding to mean and SD of each $x$.
- Calculate $y\left(x_{i}\right)$ and record the resulting values.
- The standard deviation (and distribution) of $y$ reflects the impact of $x_{i}$.

Note that the distribution of y may NOT be Gaussian, if the error propagation formula breaks down. It is then important to make this clear to the reader.

However, simulation exactly allows one to see to what degree the resulting distribution in y is Gaussian.

Error propagation exercise

## Error propagation exercise

Imagine that you measure the Length $(\mathrm{L})$ and Width $(\mathrm{W})$ of a table:

$$
\mathrm{L}=3.5 \pm 0.4 \text { and } \mathrm{W}=0.8 \pm 0.2
$$

Now you want to calculate the Perimeter (P), Area (A), and Diagonal (D), along with the uncertainty on these. This problem is an obvious case of applying the error propagation formula.

$$
\mathrm{L}=3.5 \pm 0.4
$$



But imagine that you were asked to propagate the error through:

$$
\text { y1 }=\log (\operatorname{square}(L * \tan (W))+\operatorname{sqrt}((L-W) /(\cos (W)+1.0+L)))
$$

Use simulation... not even certain that error propagation formula holds!

