

# Breaking the Spell of Gaussianity: Forecasting with higher order Fisher Matrices

Article by Elena Sellentin, Miguel Quartin and Luca Amendola. Presentation by Nikolai Plambech Nielsen

Published April 2, 2014  
Presented March 8, 2023

## Introduction

- ▶ MCMC: Robust parameter estimation, taking into account the non-Gaussianity of the system, but computationally expensive.
- ▶ Fisher Matrix: Quick parameter estimation, but approximates every likelihood as a Gaussian distribution. Includes no explicit check of non-Gaussianity.
- ▶ DALI (introduced in article): Relatively quick parameter estimation, takes into account non-Gaussianities.

## General likelihood form

Authors assume a general Gaussian form of the likelihood:

$$L = N \exp \left[ -\frac{1}{2} [\mathbf{m} - \boldsymbol{\mu}(\mathbf{p})]^T M [\mathbf{m} - \boldsymbol{\mu}(\mathbf{p})] \right] \quad (1)$$

where  $\mathbf{m}$  is vector of measurements,  $\boldsymbol{\mu}(\mathbf{p})$  is theoretical predictions corresponding to the measurements (evaluated at the parameters  $\mathbf{p}$ ), and  $M$  is the inverse covariance matrix.

Note: Parametric dependence is only in the model  $\mu$ . In appendix and future work the authors generalize to covariances with parametric dependence.

## Fisher Matrix

Negative Hessian of LLH. In frequentist regime this reduces to

$$F_{\alpha\beta} = \langle \mathcal{L}_{,\alpha} \mathcal{L}_{,\beta} \rangle, \quad \mathcal{L} = \ln(L). \quad (2)$$

LLH is

$$\mathcal{L} \approx N - \frac{1}{2} \boldsymbol{\mu}_{,\alpha} M \boldsymbol{\mu}_{,\beta} \Delta p_{\alpha} \Delta p_{\beta} \equiv N + F, \quad (3)$$

where,  $\Delta p_{\alpha} = p_{\alpha} - \hat{p}_{\alpha}$  is deviation from MLE parameters. Note parabolic in parameters, so always Gaussian.

---

<sup>0</sup>Subscript ", $\alpha$ " means a partial derivative w.r.t.  $\alpha$ .

## DALI: Taylor expansion of likelihood

To fourth order in parameters:

$$\begin{aligned}\mathcal{L} \approx N & - \frac{1}{2} F_{\alpha\beta} \Delta p_{\alpha} \Delta p_{\beta} \\ & - \frac{1}{3!} S_{\alpha\beta\gamma} \Delta p_{\alpha} \Delta p_{\beta} \Delta p_{\gamma} \\ & - \frac{1}{4!} Q_{\alpha\beta\gamma\delta} \Delta p_{\alpha} \Delta p_{\beta} \Delta p_{\gamma} \Delta p_{\delta},\end{aligned}\tag{4}$$

with

$$F_{\alpha\beta} = \mathcal{L}_{,\alpha\beta}, \quad S_{\alpha\beta\gamma} = \mathcal{L}_{,\alpha\beta\gamma}, \quad Q_{\alpha\beta\gamma\delta} = \mathcal{L}_{,\alpha\beta\gamma\delta}\tag{5}$$

Problem:  $Q$ - and  $S$ -terms not globally negative, and thus non-normalizable.

## DALI: Collecting terms in orders of derivatives - Doublet DALI

To second order in the derivatives of the model we have the  
"doublet DALI"

$$\begin{aligned}\mathcal{L} &\approx N + F - \left[ \frac{1}{2} \boldsymbol{\mu}_{,\alpha\beta} M \boldsymbol{\mu}_{,\gamma} \Delta p_\alpha \Delta p_\beta \Delta p_\gamma \right. \\ &\quad \left. + \frac{1}{8} \boldsymbol{\mu}_{,\delta\gamma} M \boldsymbol{\mu}_{,\beta\alpha} \Delta p_\alpha \Delta p_\beta \Delta p_\gamma \Delta p_\delta \right] \quad (6) \\ &= N + F + S\end{aligned}$$

Leading second order term at big  $p$  is now negative, if  $M$  is positive definite, so entire approximation is normalizable.

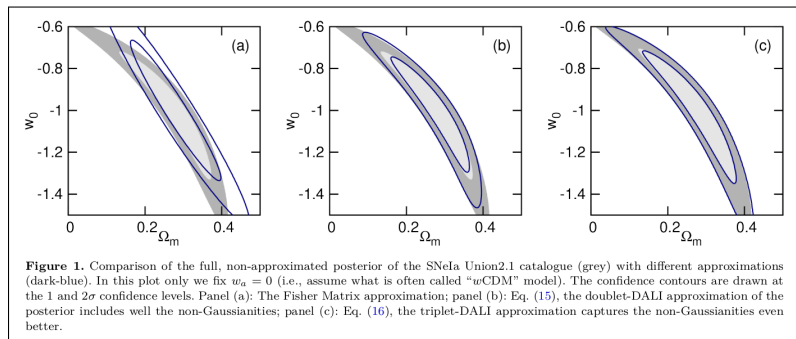
## DALI: Collecting terms in orders of derivatives - Triplet DALI

To third order, the "triplet DALI":

$$\begin{aligned}\mathcal{L} \approx N + F + S - & \left[ \frac{1}{6} \boldsymbol{\mu}_{,\delta} M \boldsymbol{\mu}_{,\beta\alpha\gamma} \Delta p_\alpha \Delta p_\beta \Delta p_\gamma \Delta p_\delta \right. \\ & + \frac{1}{12} \boldsymbol{\mu}_{,\alpha\beta\delta} M \boldsymbol{\mu}_{,\gamma\tau} \Delta p_\alpha \Delta p_\beta \Delta p_\gamma \Delta p_\delta \Delta p_\tau \\ & \left. + \frac{1}{72} \boldsymbol{\mu}_{,\alpha\beta\delta} M \boldsymbol{\mu}_{,\delta\tau\sigma} \Delta p_\alpha \Delta p_\beta \Delta p_\gamma \Delta p_\delta \Delta p_\tau \Delta p_\sigma \right] \\ = N + F + S + Q.\end{aligned}\tag{7}$$

# Applications

Applied to type Ia supernova data (Union2.1 catalogue)

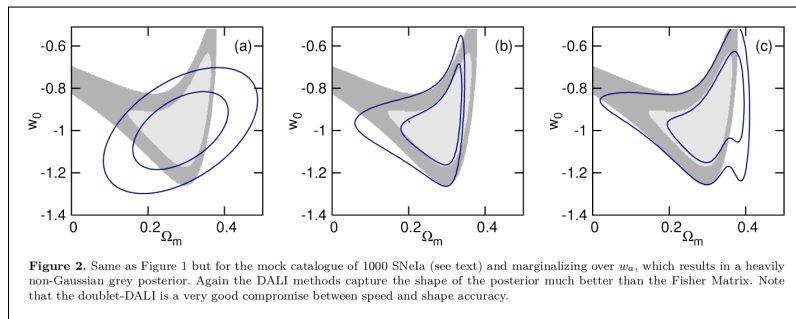


**Figure:** The DALI method applied to mock SNeIa data, from the article. The LLH is calculated as a function of the mass-density  $\Omega_m$  and the present dark energy equation of state parameter  $w_0$ .



# Applications

Applied to mock type Ia supernova data



**Figure:** The DALI method applied to SNeIa data from the Union2.1 catalogue, from the article. The LLH is calculated as a function of the mass-density  $\Omega_m$  and the present dark energy equation of state parameter  $w_0$ . Here marginalized over  $w_a \in (-\infty, \infty)$ , from the "CPL" model.