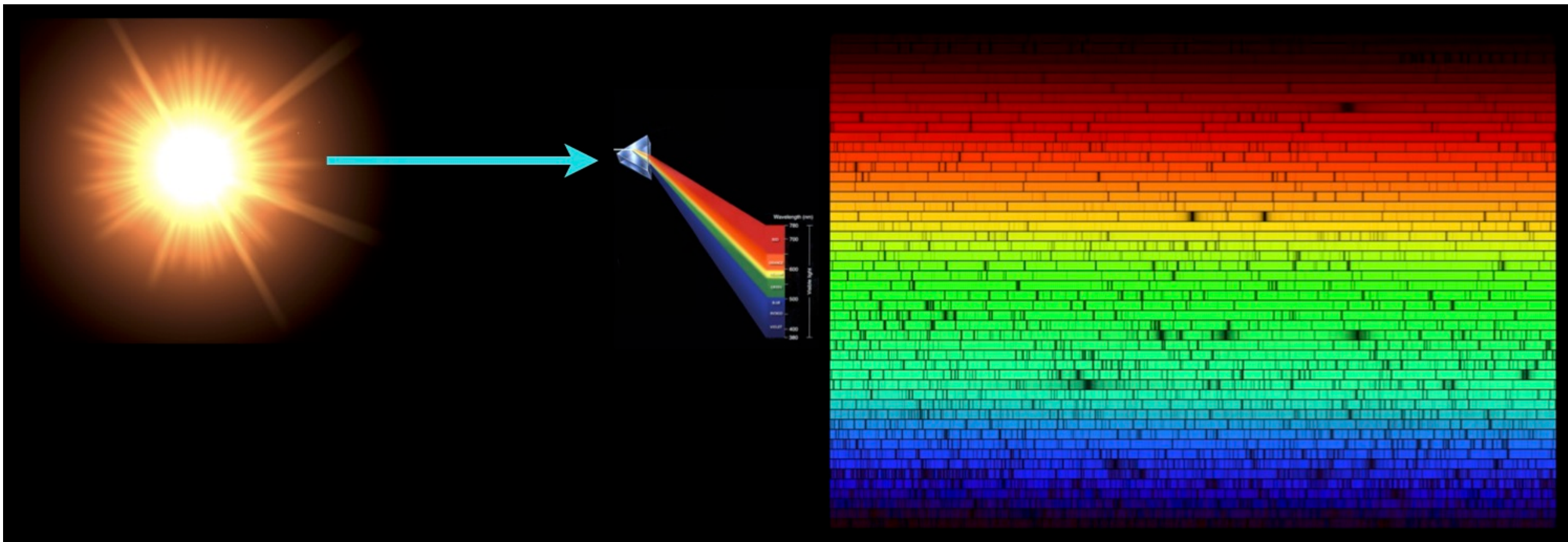


Stellar Classification with Neural Networks

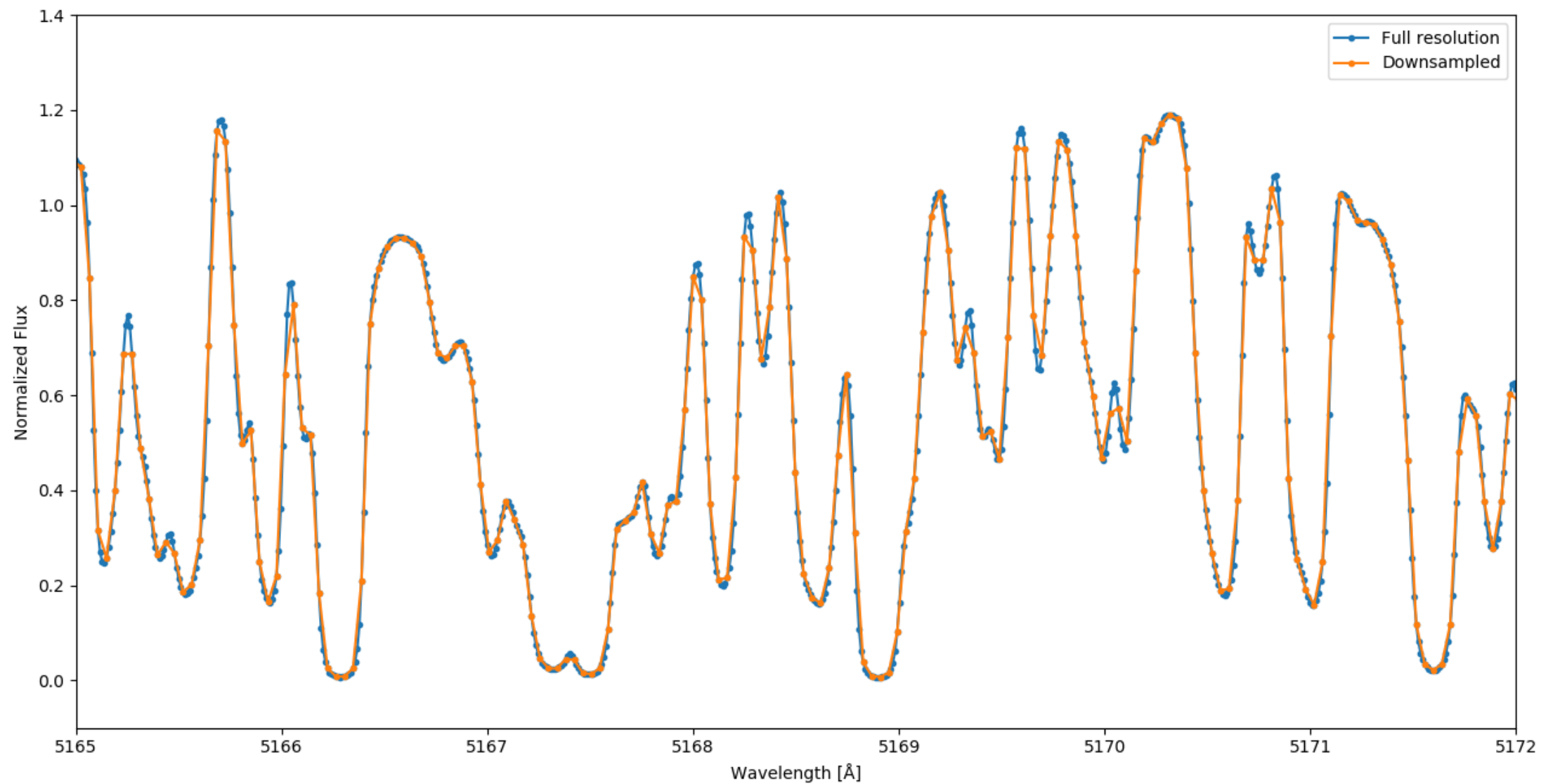
Zlatko Saldic & Alexander Rathcke

(All group members have contributed evenly to the project)

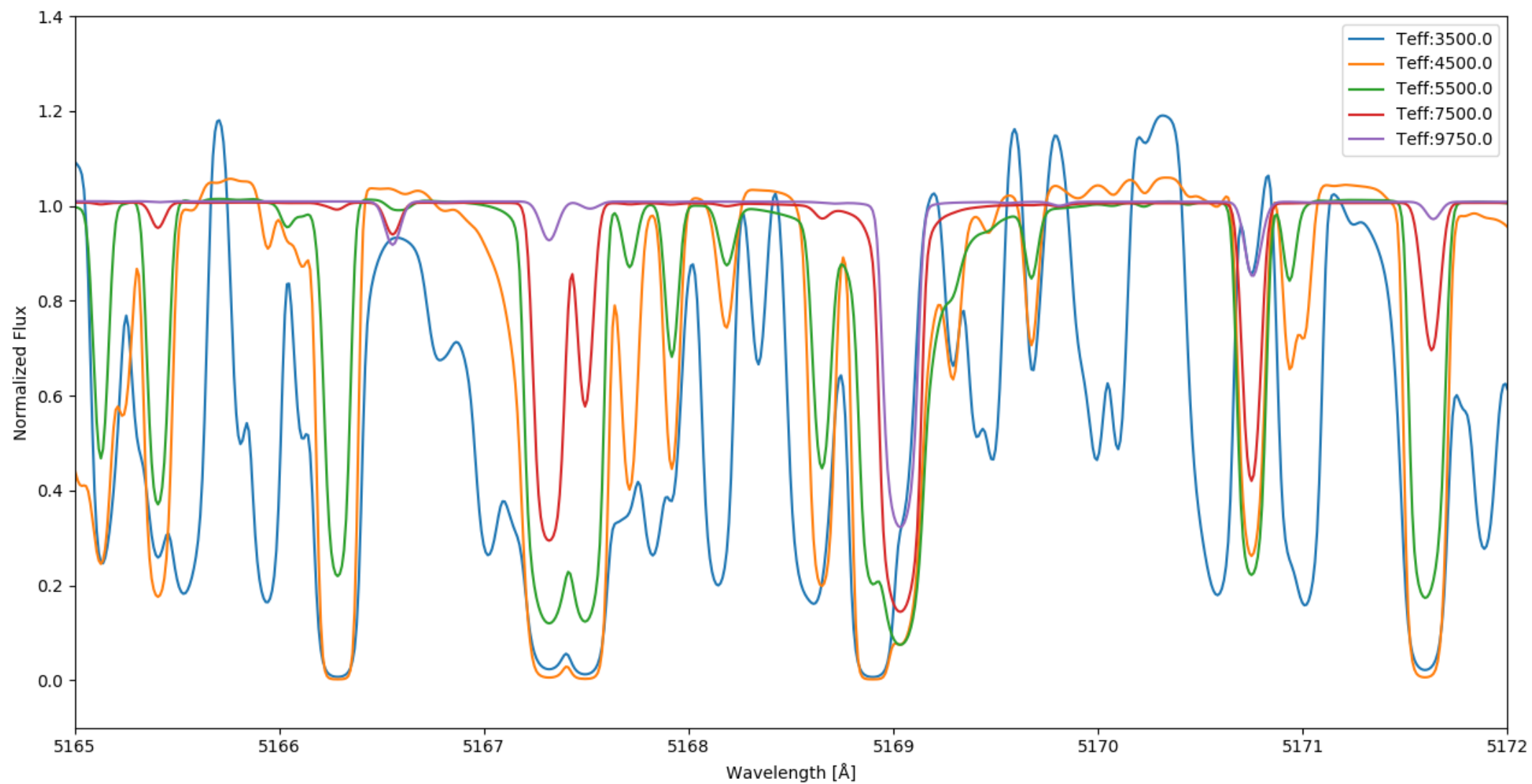


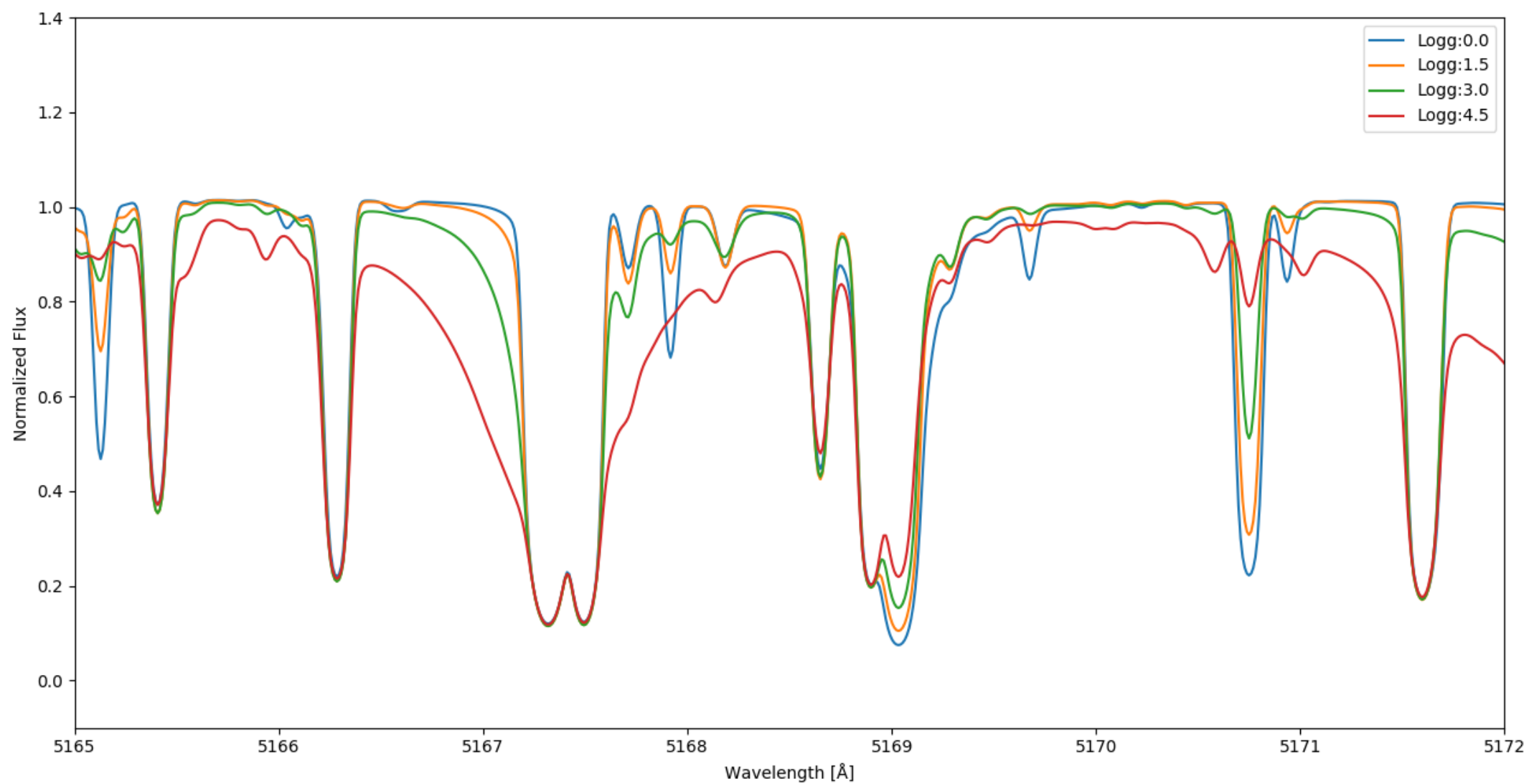
Data:

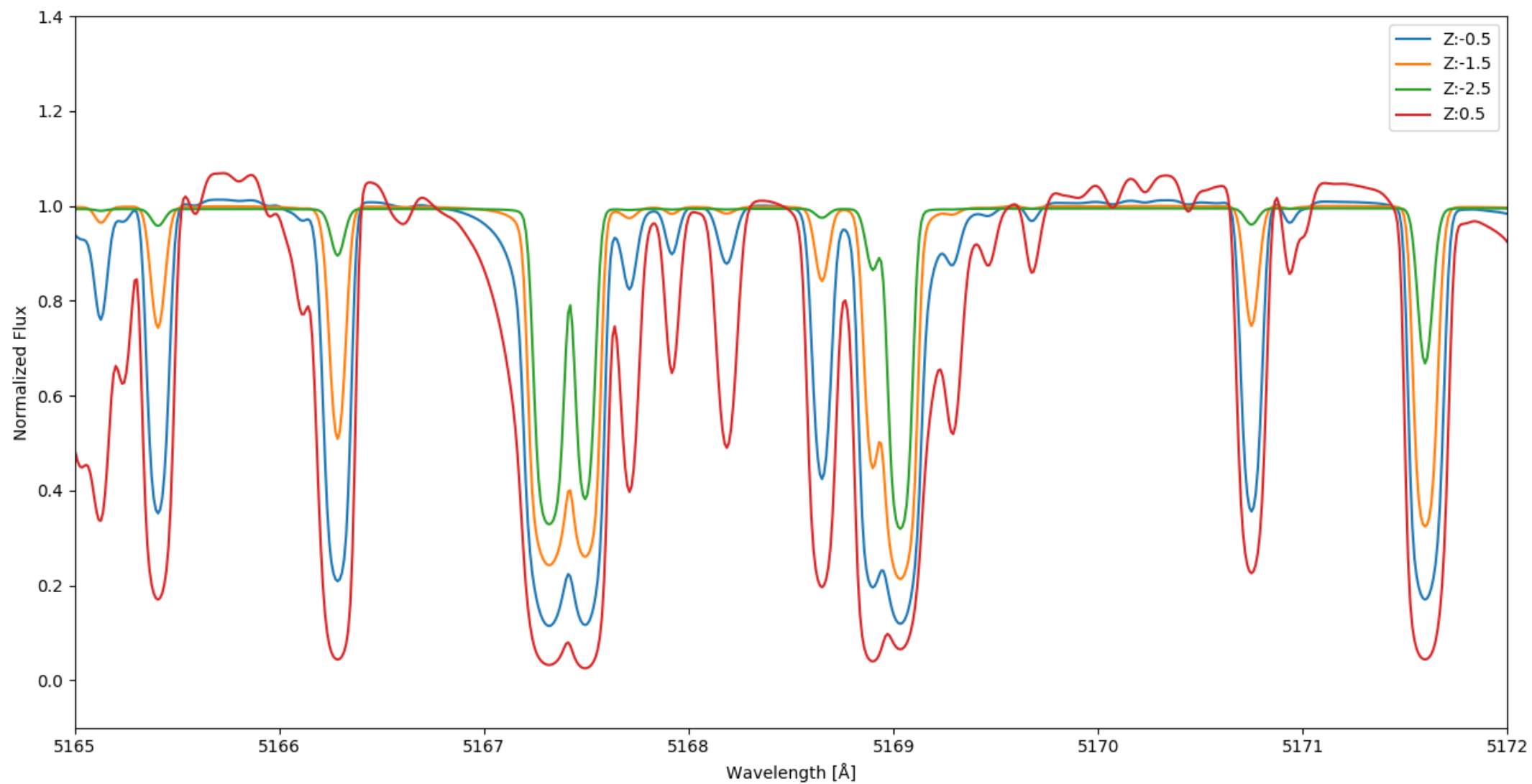
- Model grid consisting of ~50k synthetic stellar spectra
- ~30k pixels covering a spectral range of 300 Å (very high resolution)
- 4 target parameters (labeled)
- Preprocessing:
 - Normalizing spectrum and labels
 - Spectrum downsampling (X.shape: (51359, 29788) -> X.shape(51359, 7447))

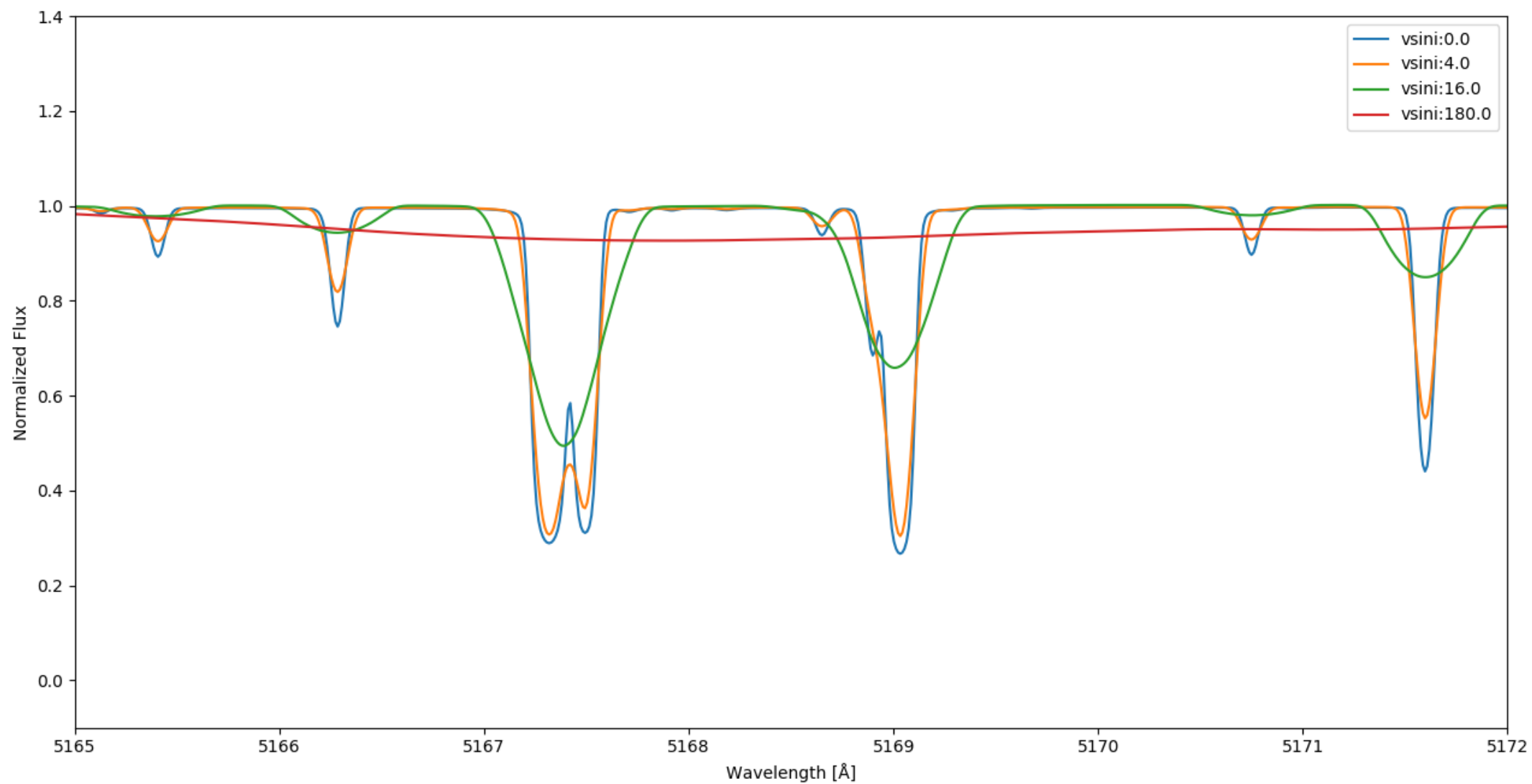


[illegible]









Approaching the problem & building the network

Similar problems exists

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Introduction to 1D Convolutional Neural Networks in Keras for Time Sequences



Nils Ackermann

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Sep 4, 2018 · 7 min read

Introduction

Many articles focus on two dimensional conv.

When to Apply a 1D CNN?

A CNN works well for identifying simple patterns within your data which will then be used to form more complex patterns within higher layers. A 1D CNN

is very effective when you expect to derive interesting features from shorter (fixed-length) segments of the overall data set and where the location of the feature within the segment is not of high relevance.

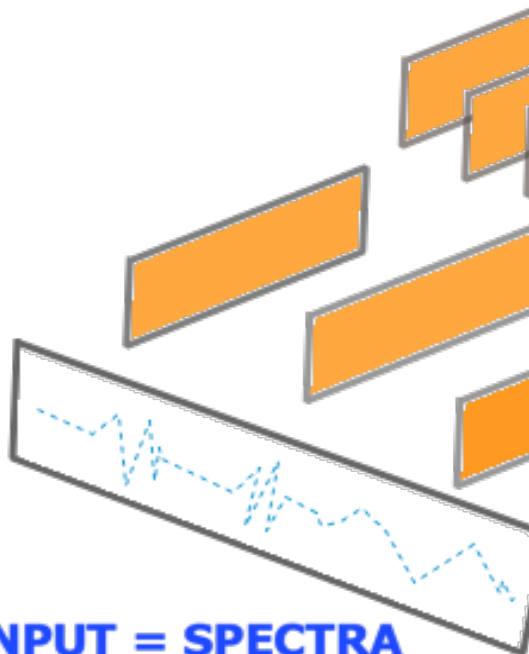
Initial structure for final network adopted from:
Fabbro, Sebastien, et al. (2017)

OUTPUT = TEMPERATURE

The Neural

**FULLY CONNECT
LAYERS**

MAX POOLING



INPUT = SPECTRA

Layer (type)	Output Shape	Param #
Input (InputLayer)	(None, 7447)	0
reshape_1 (Reshape)	(None, 7447, 1)	0
conv1d_1 (Conv1D)	(None, 7447, 4)	132
conv1d_2 (Conv1D)	(None, 7447, 16)	2064
max_pooling1d_1 (MaxPooling1	(None, 1861, 16)	0
flatten_1 (Flatten)	(None, 29776)	0
dense_1 (Dense)	(None, 256)	7622912
dense_2 (Dense)	(None, 128)	32896
Output (Dense)	(None, 4)	516
Total params: 7,658,520		
Trainable params: 7,658,520		
Non-trainable params: 0		

Some nice KERAS features

```
38016/38519 [=====>.] - ETA: 3s - loss: 0.0013 - mean_absolute_error: 0.0242
38048/38519 [=====>.] - ETA: 3s - loss: 0.0013 - mean_absolute_error: 0.0242
38080/38519 [=====>.] - ETA: 3s - loss: 0.0013 - mean_absolute_error: 0.0242
38112/38519 [=====>.] - ETA: 2s - loss: 0.0013 - mean_absolute_error: 0.0242
38144/38519 [=====>.] - ETA: 2s - loss: 0.0013 - mean_absolute_error: 0.0242
38176/38519 [=====>.] - ETA: 2s - loss: 0.0013 - mean_absolute_error: 0.0242
38208/38519 [=====>.] - ETA: 2s - loss: 0.0013 - mean_absolute_error: 0.0242
38240/38519 [=====>.] - ETA: 1s - loss: 0.0013 - mean_absolute_error: 0.0242
38272/38519 [=====>.] - ETA: 1s - loss: 0.0013 - mean_absolute_error: 0.0242
38304/38519 [=====>.] - ETA: 1s - loss: 0.0013 - mean_absolute_error: 0.0242
38336/38519 [=====>.] - ETA: 1s - loss: 0.0013 - mean_absolute_error: 0.0242
38368/38519 [=====>.] - ETA: 1s - loss: 0.0013 - mean_absolute_error: 0.0242
38400/38519 [=====>.] - ETA: 0s - loss: 0.0013 - mean_absolute_error: 0.0242
38432/38519 [=====>.] - ETA: 0s - loss: 0.0013 - mean_absolute_error: 0.0242
38464/38519 [=====>.] - ETA: 0s - loss: 0.0013 - mean_absolute_error: 0.0242
38496/38519 [=====>.] - ETA: 0s - loss: 0.0013 - mean_absolute_error: 0.0242
38519/38519 [=====] - 287s 7ms/step - loss: 0.0013 - mean_absolute_error: 0.0242 -
val_loss: 0.0014 - val_mean_absolute_error: 0.0245
Epoch 00023: early stopping
trained_data.h5 saved.

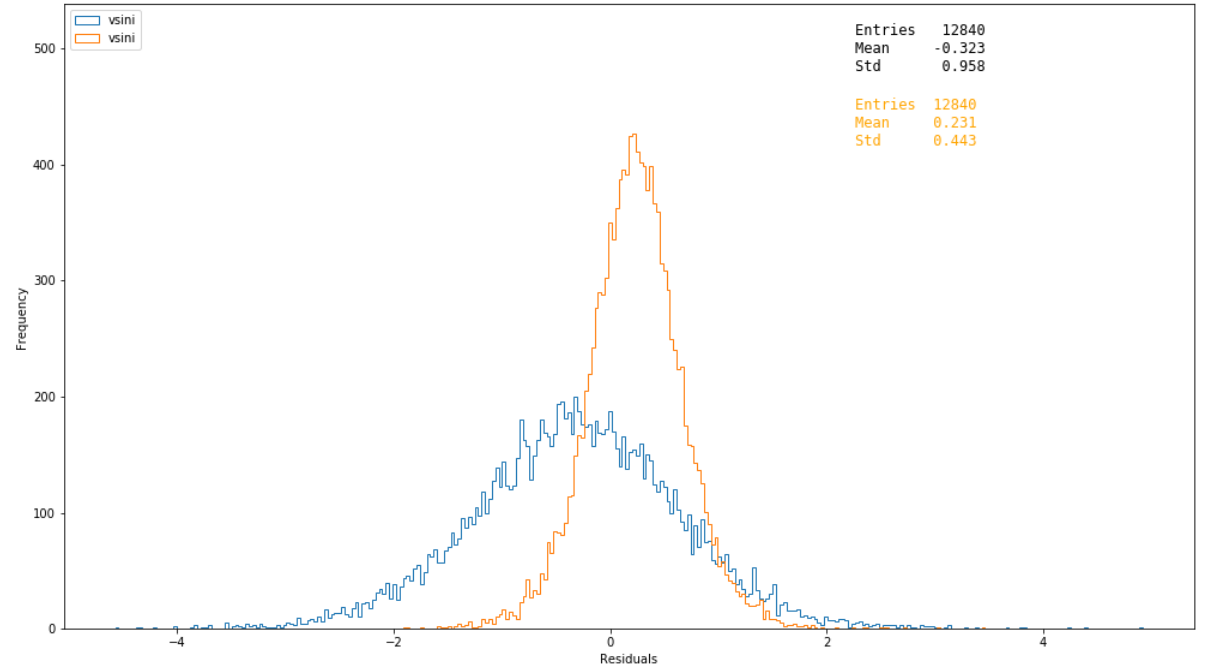
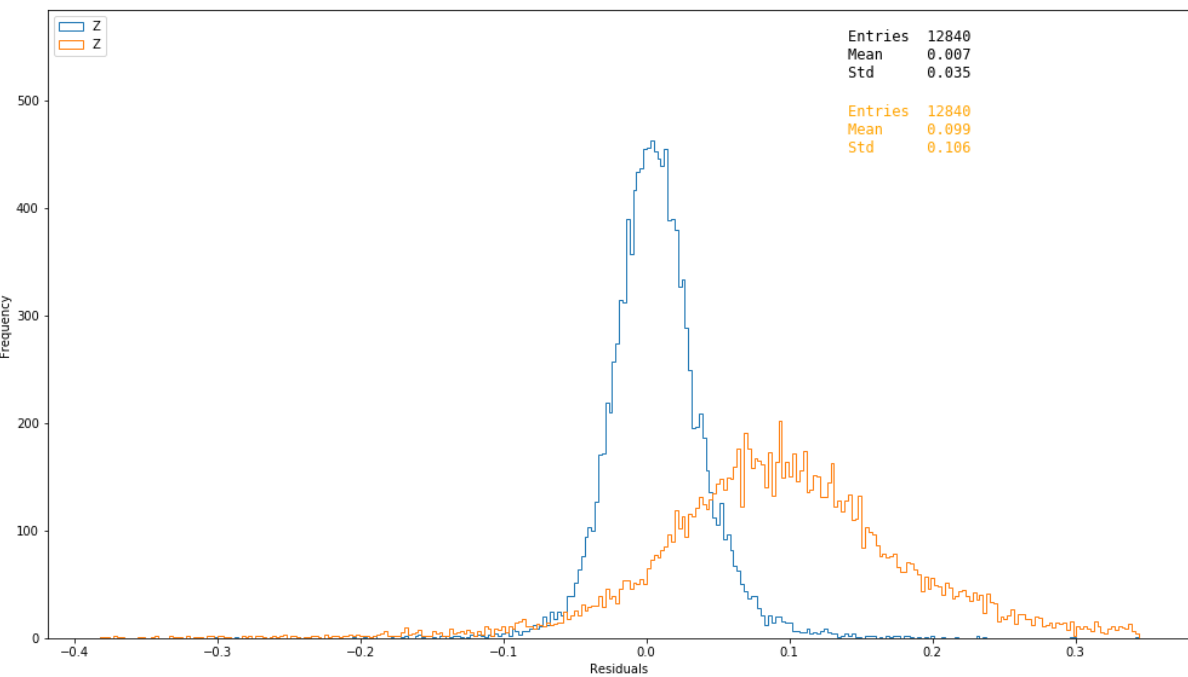
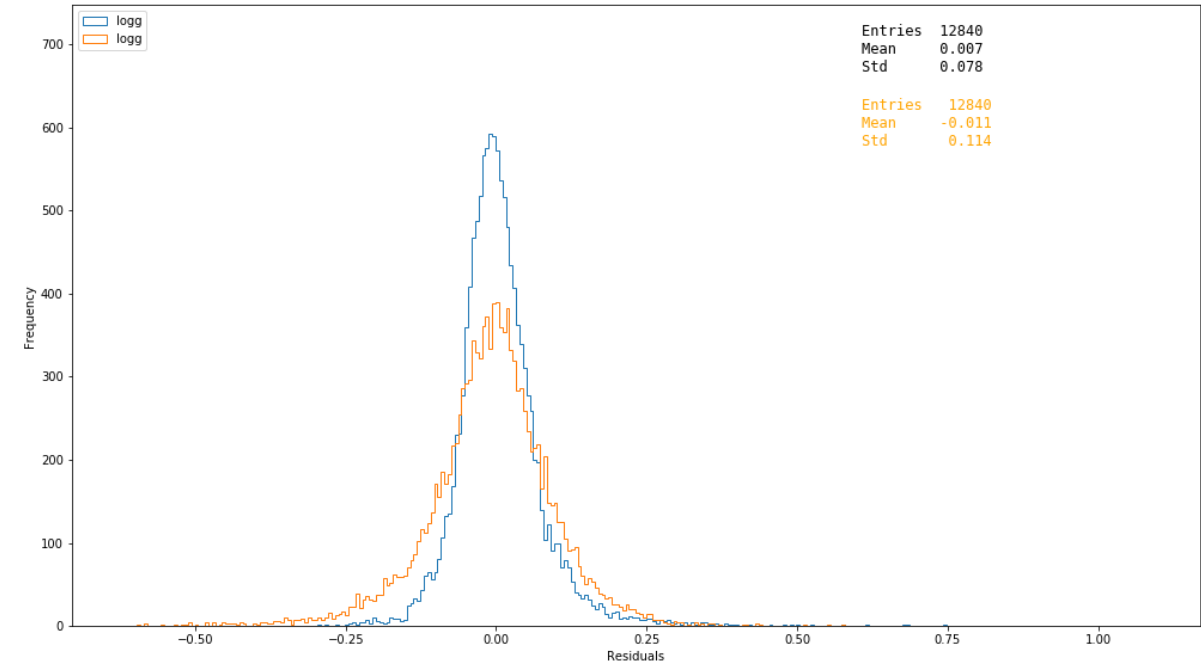
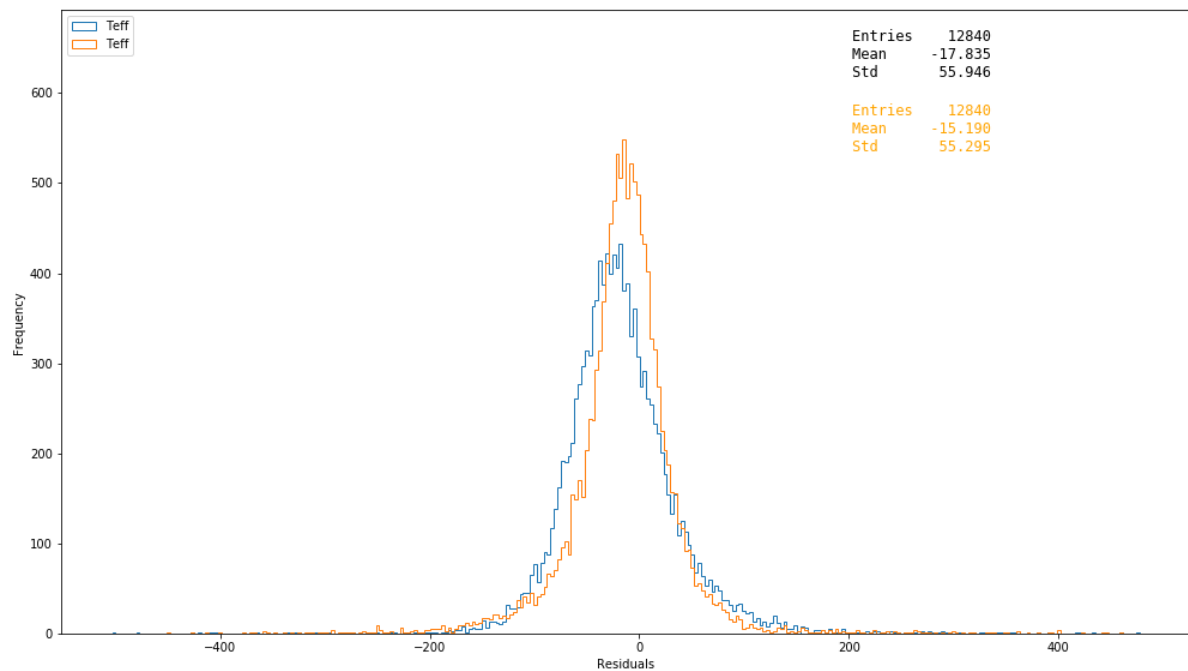
Process finished with exit code 0
```

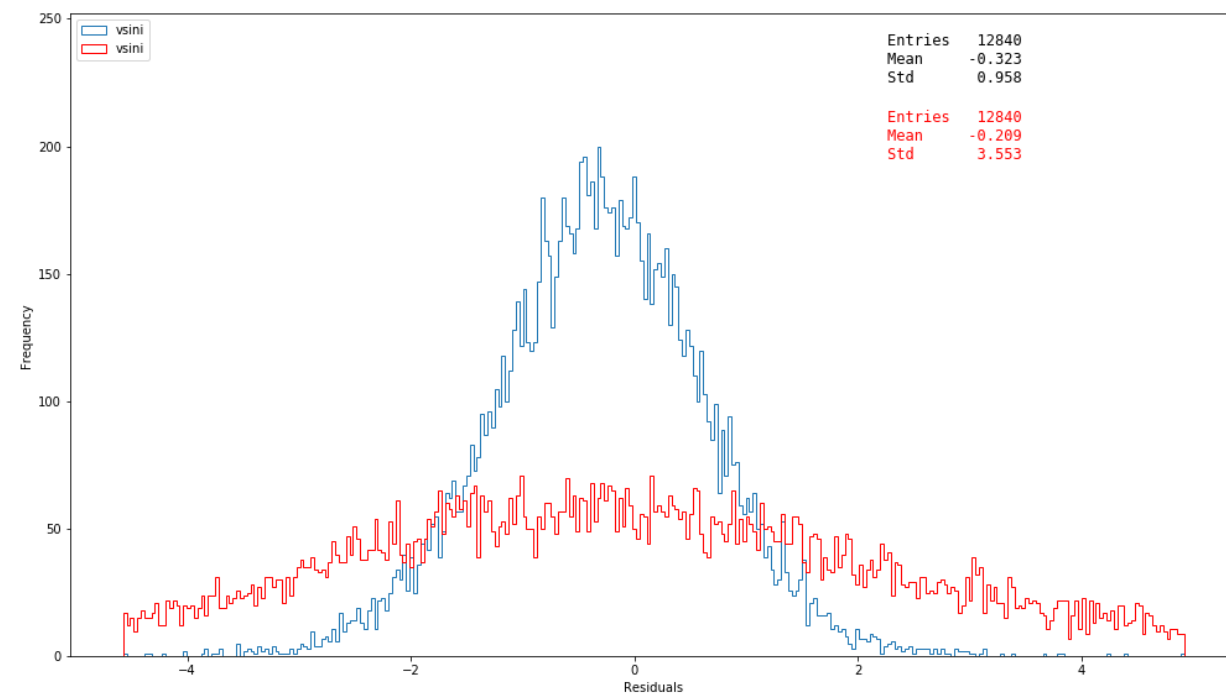
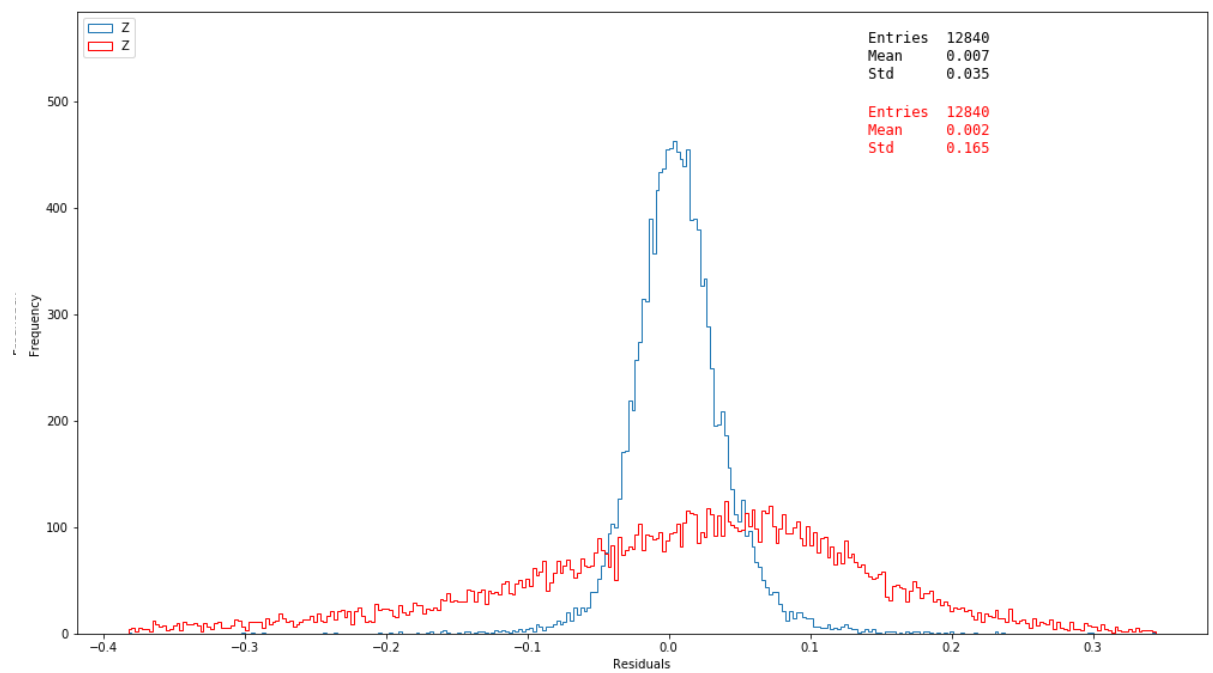
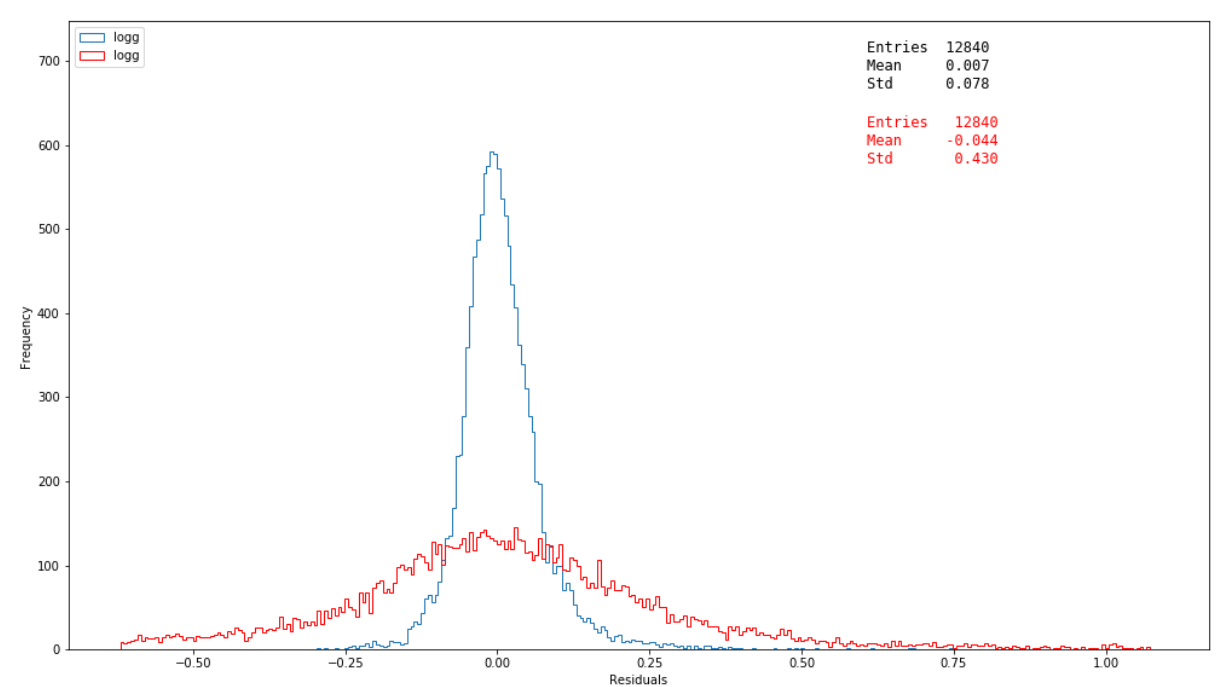
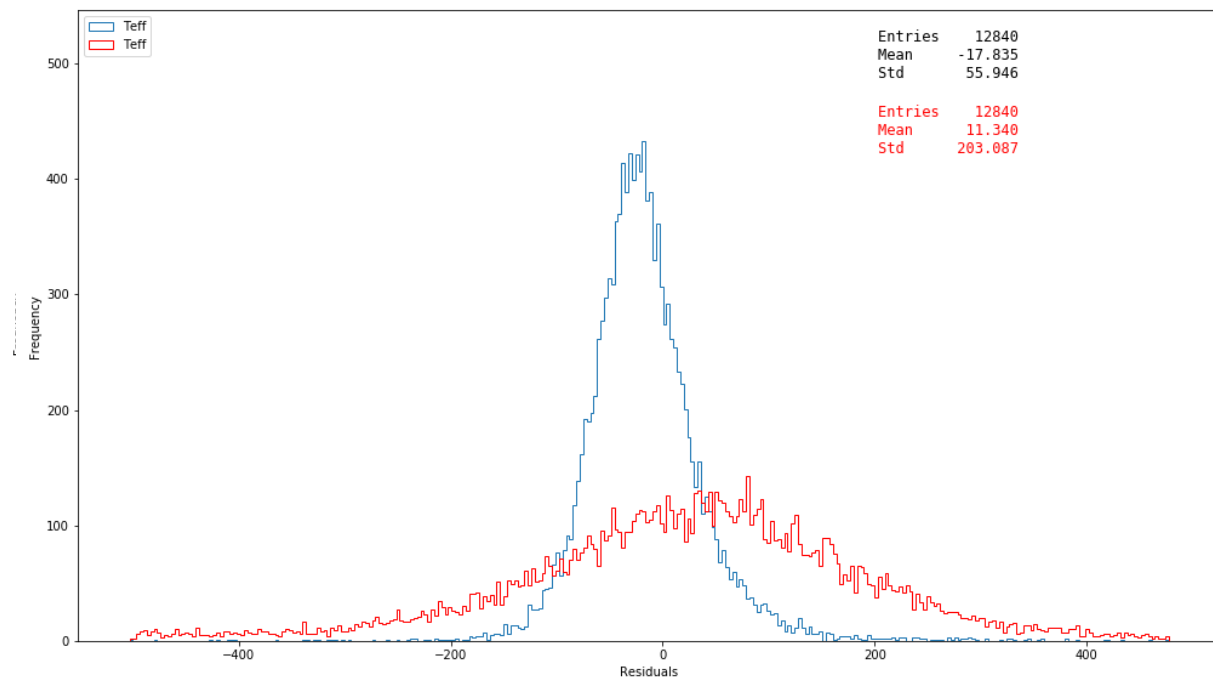
'min')

Hyper parameters

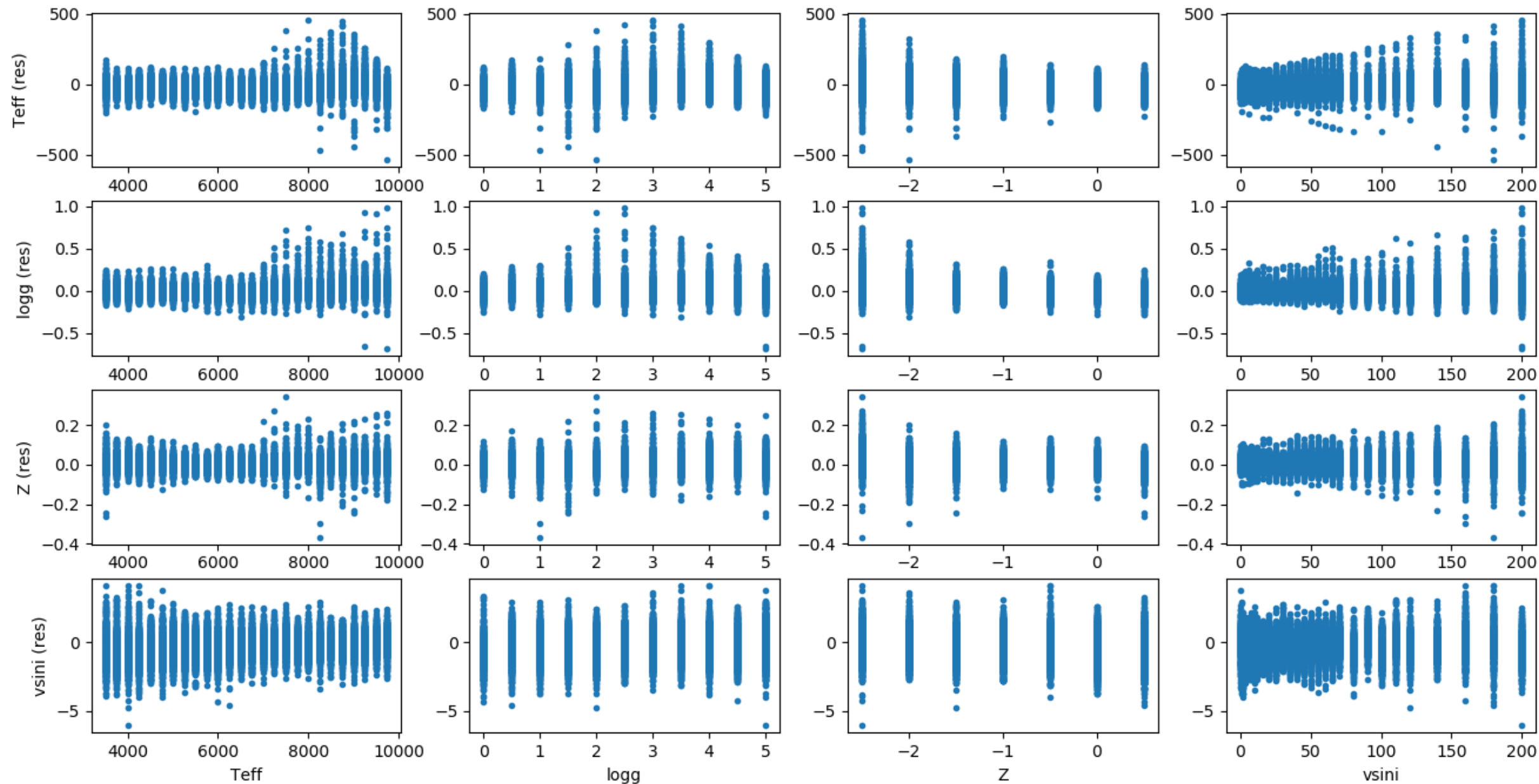
- Kernel size (filter length that convolves across spectrum) = 32 pixels
w/ 32 pixels we were sure about detecting important features in the spectrum
- Activation = ReLu
- Gradient Descent Optimizer = ADAM
- Maxpool window length (the number of inputs which are compared against each other to find the max value) = 4
- Learning rate (of the gradient descent optimizer) = 0.0007
- Batch size (number of spectra fed into the model at once) = 64
- Loss function = Mean Squared Error

Results





For multivariate regression approach:



Perspective & Future Work

- Make applicable to real observations
- Hyperparameter optimization using GPU cluster
- Expand to estimate alpha particle enrichment (or even individual abundances)
- Error estimation (Bayesian Neural Networks?)

Questions?