

Inferring Supernova dust with Neural Networks

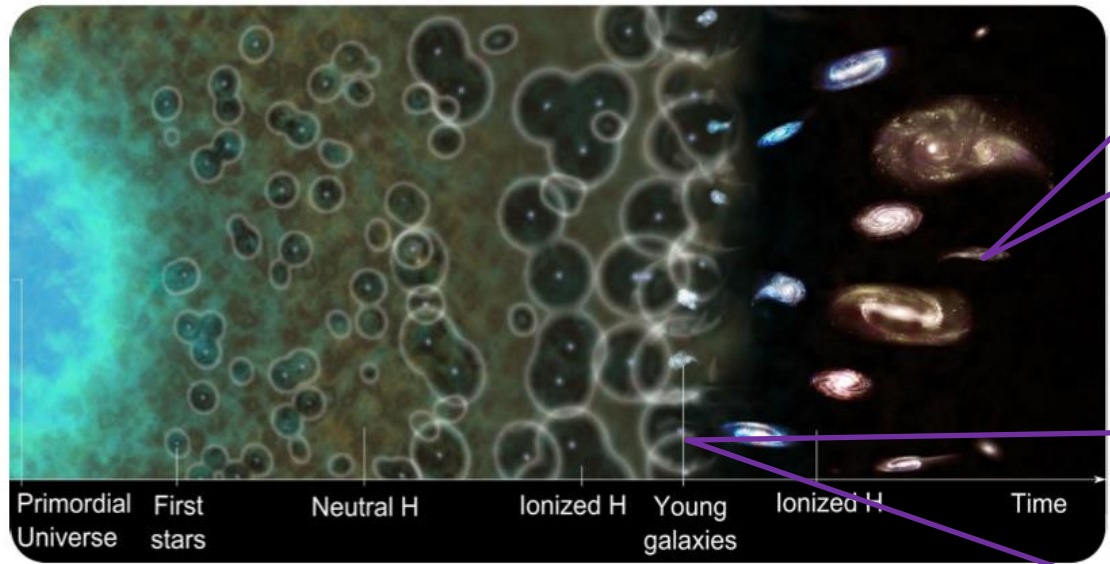
Zoe Ansari – May 2021



- What is cosmic dust
- Why Machine Learning?
- Mixture Density Network
- Uncertainty
- Recursive feature elimination

Why should we keep an eye on dust

30 ← Epoch of Reionization → 6 redshift
0.10 ← 0.93 time (Gyrs)



$Z \sim 0.0001$

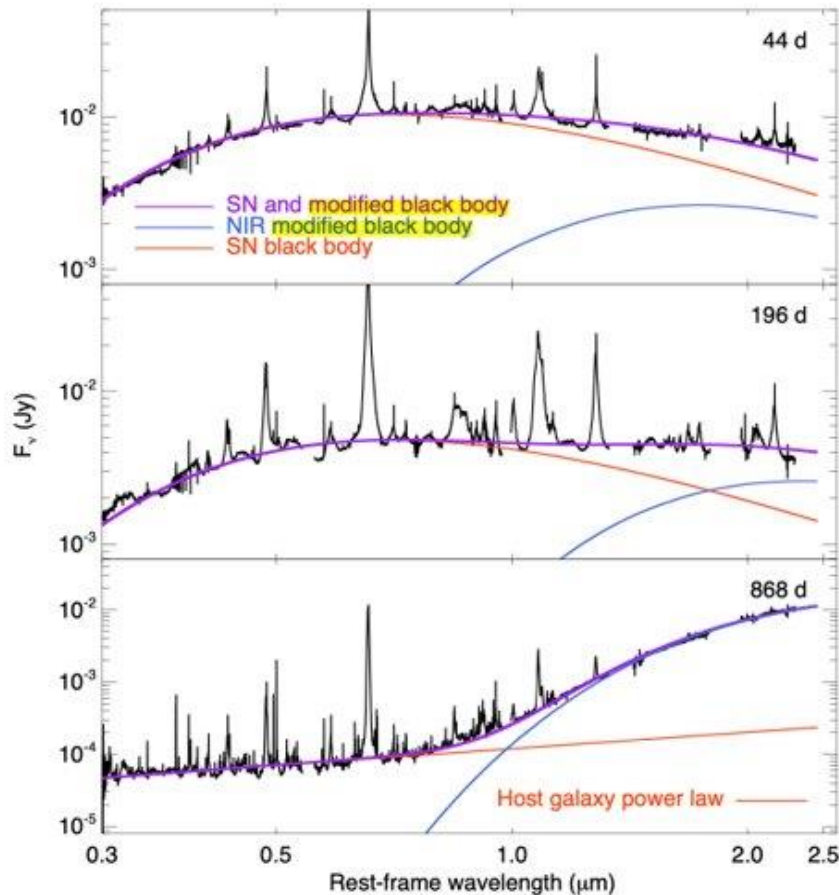
Supernovae : CCSNe
Consistency in
observations
(Gall & Hjorth 2018)

$Z \sim 6-7$

http://www.oa.uj.edu.pl/cm2014/talks/durrive_jb.pdf

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Challenges in inferring dust from observations



What we detect - spectra

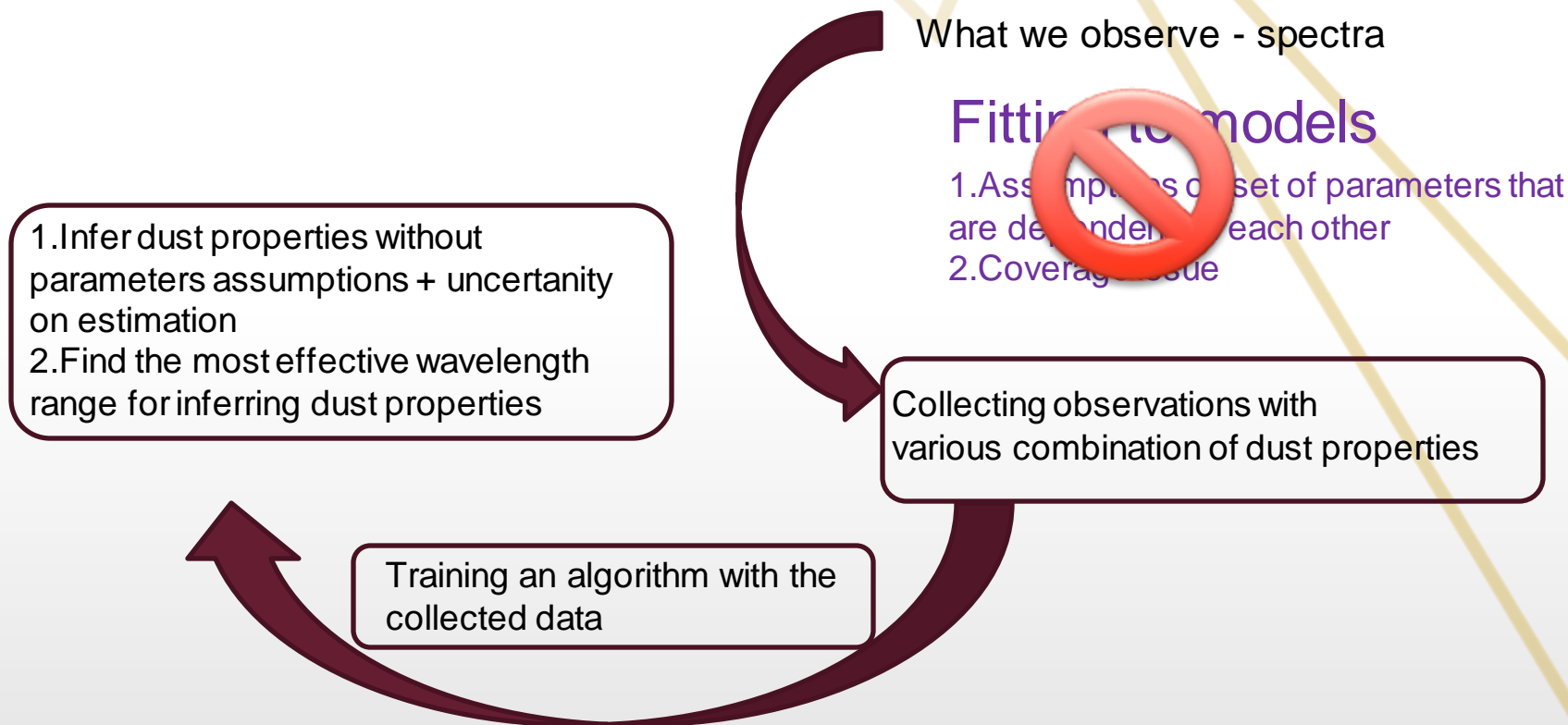
Fitting to models

1. Assumptions on set of parameters that are dependent on each other
2. Coverage issue



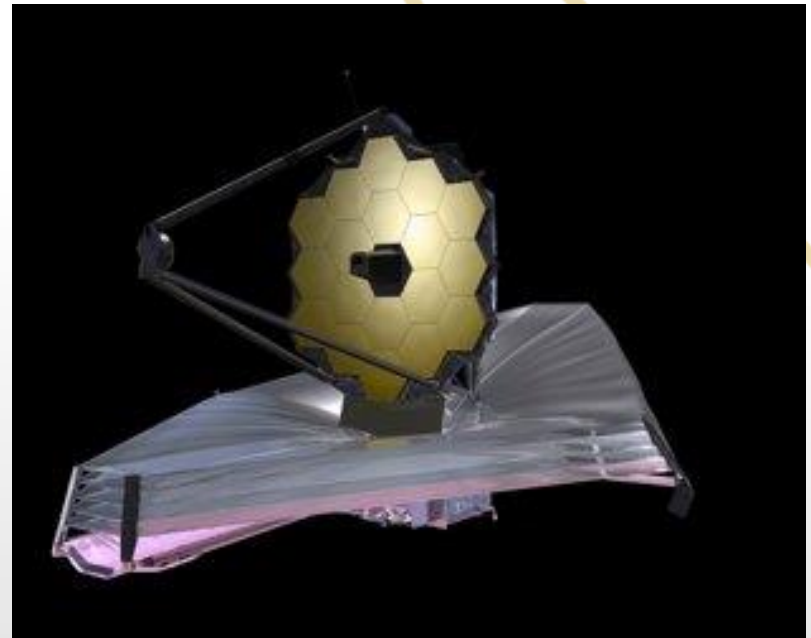
Dust mass
Dust temp.
Dust species

Inferring dust from observations with Machine



Collect the data

James Webb Space Telescope

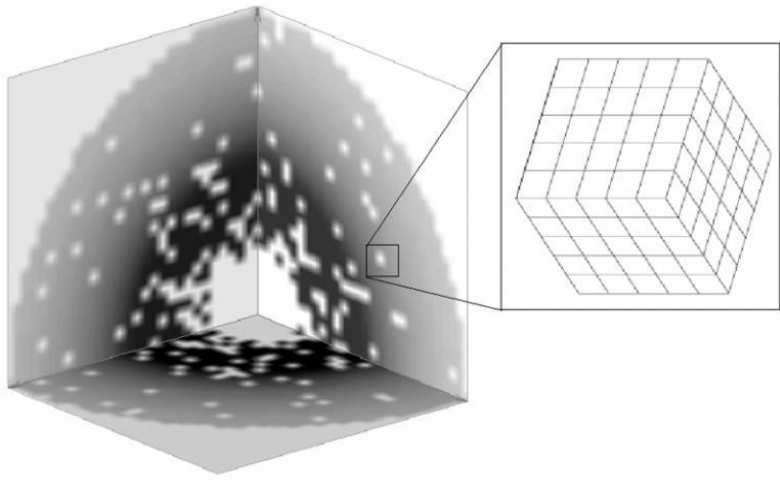


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Collect the data

Simulations

MOCASSIN



3D representative of supernova
ejecta model

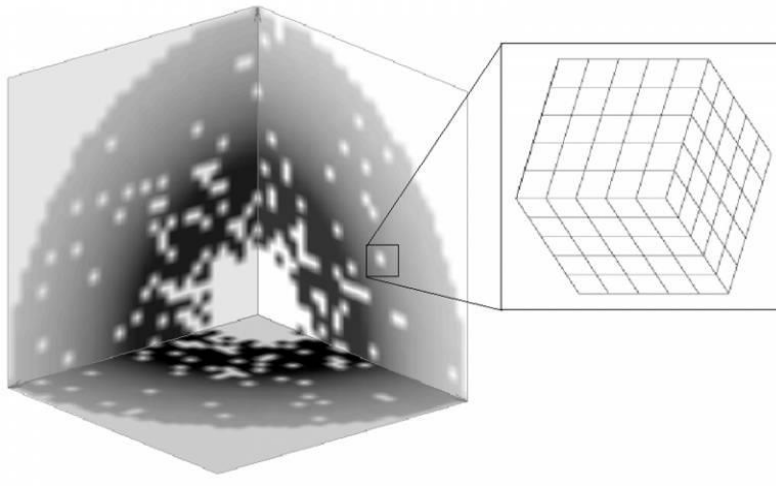
James Webb Space Telescope



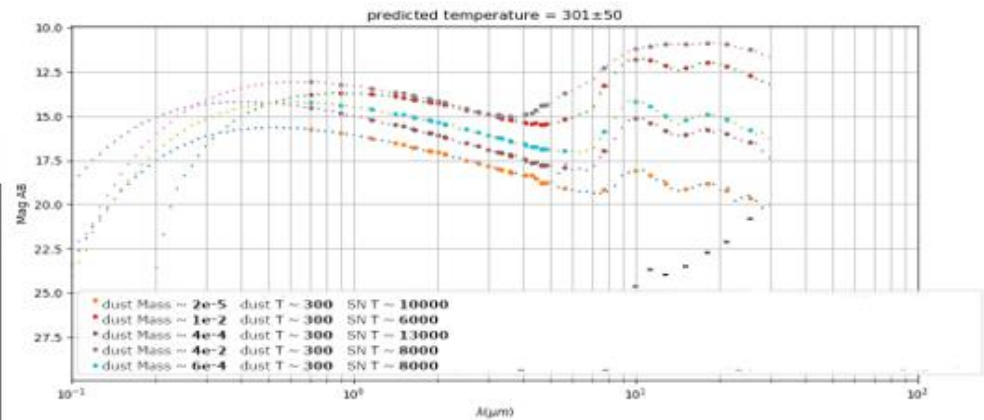
Collect the data

Simulations

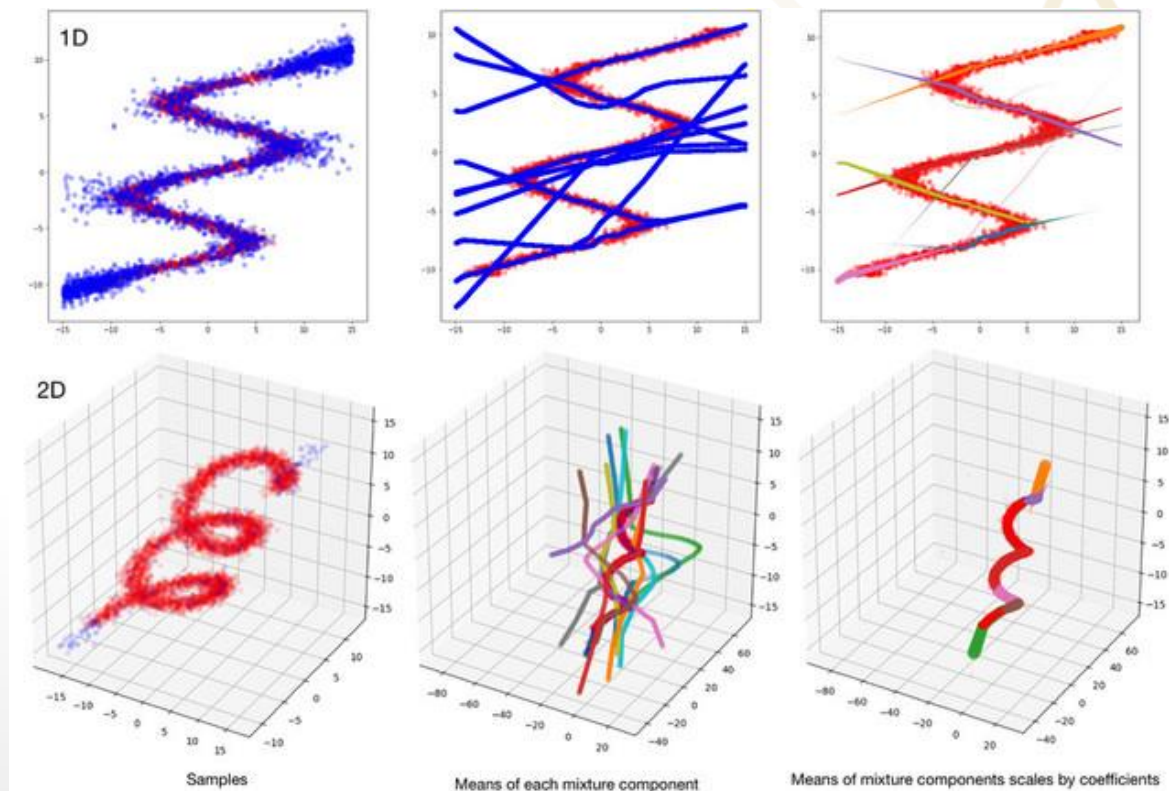
MOCASSIN



3D representative of supernova
ejecta model



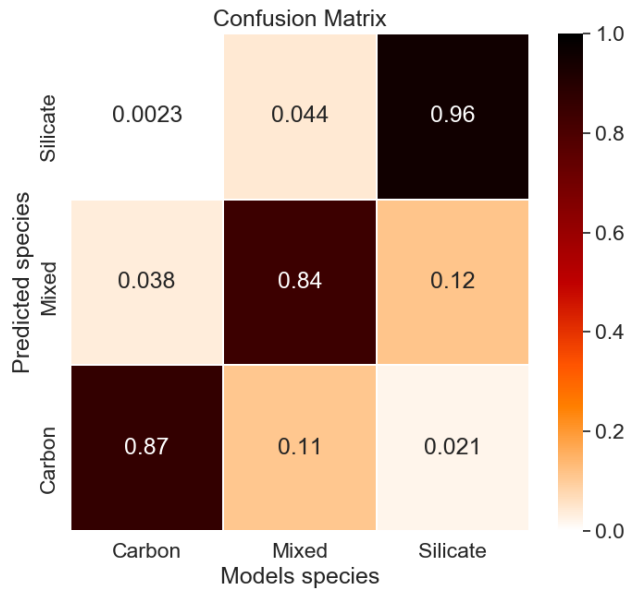
Mixture Density Network



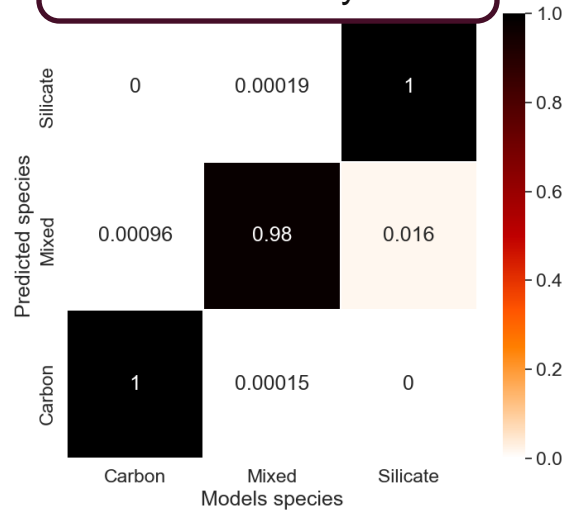
$$E(\mathbf{w}) = -\sum_{n=1}^N \ln \left\{ \sum_{k=1}^K \pi_k(\mathbf{x}_n, \mathbf{w}) N(\mathbf{t} | \mu_k(\mathbf{x}_n, \mathbf{w}), \sigma_k^2(\mathbf{x}_n, \mathbf{w})) \right\}$$

<https://github.com/cpmpercussion/keras-mdn-layer>

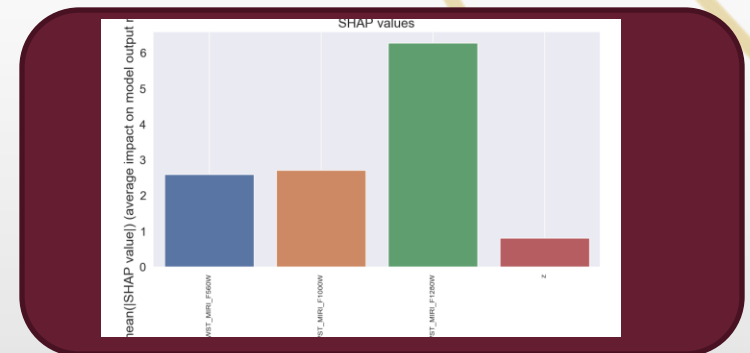
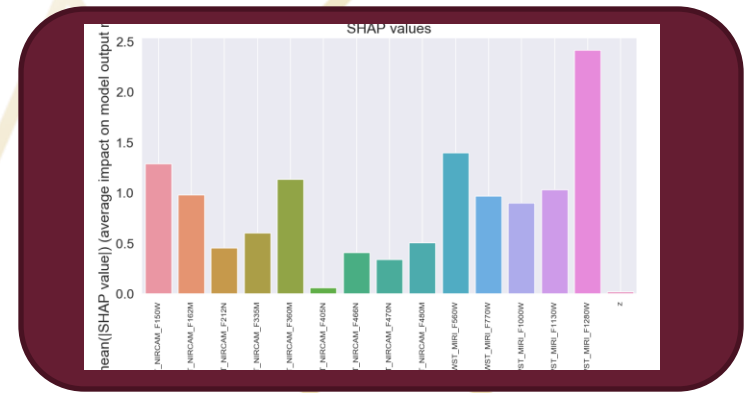
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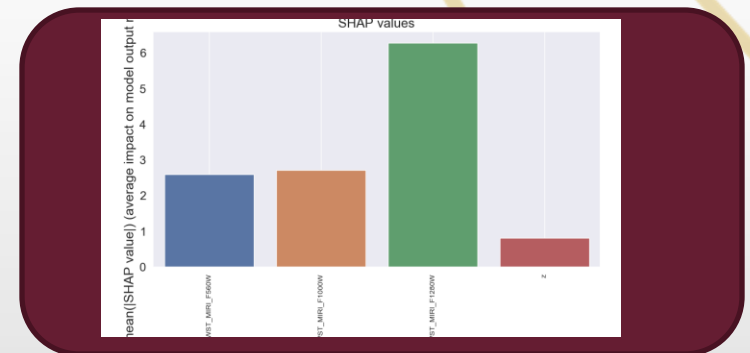
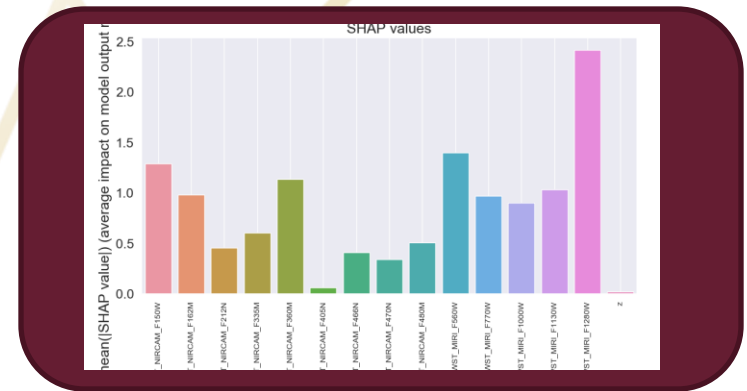
Estimates with low uncertainty



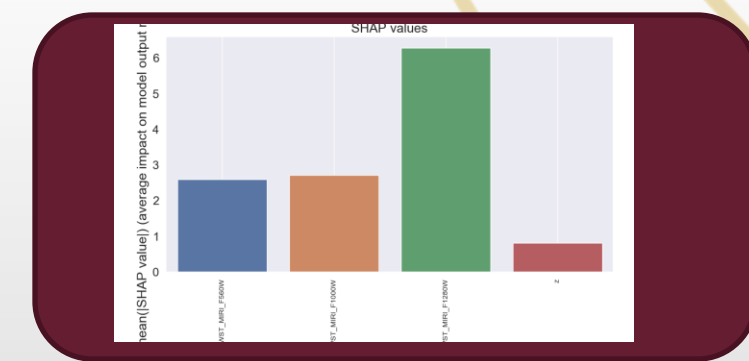
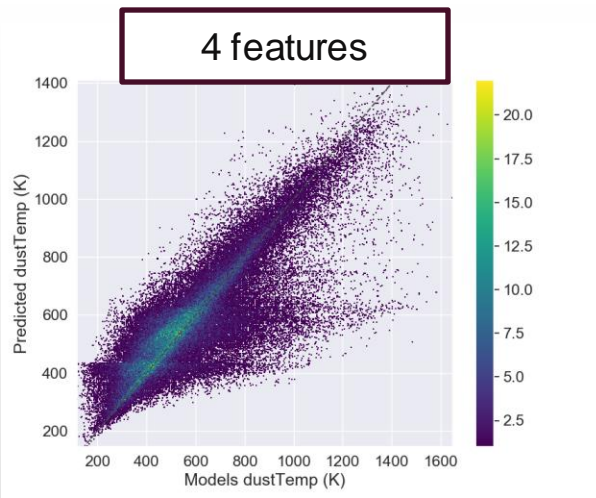
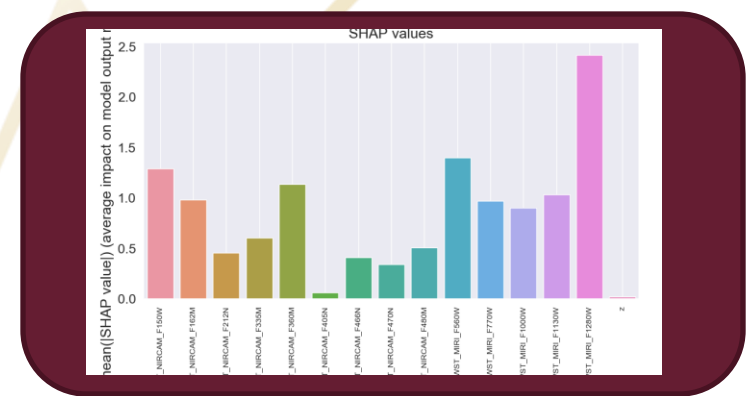
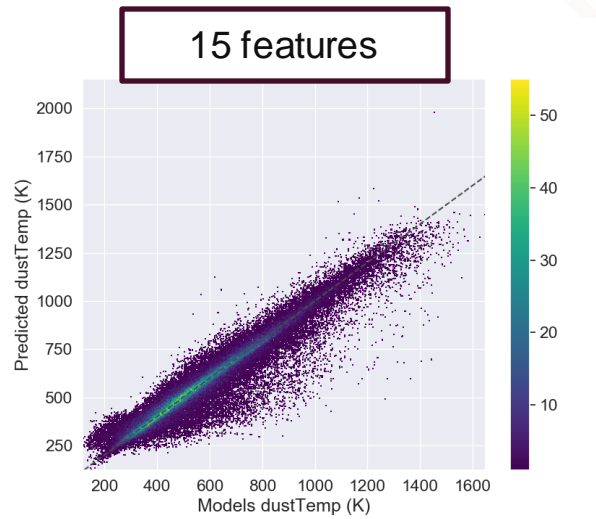
Recursive feature elimination



Recursive feature elimination



Recursive feature elimination





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