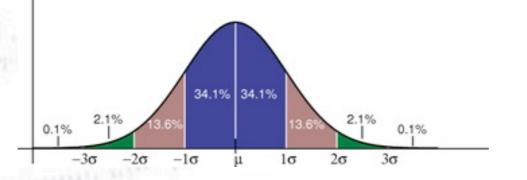
# Applied Statistics Confidence intervals and Limits



Troels C. Petersen (NBI)

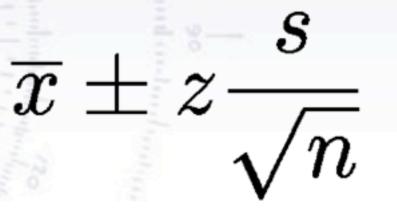


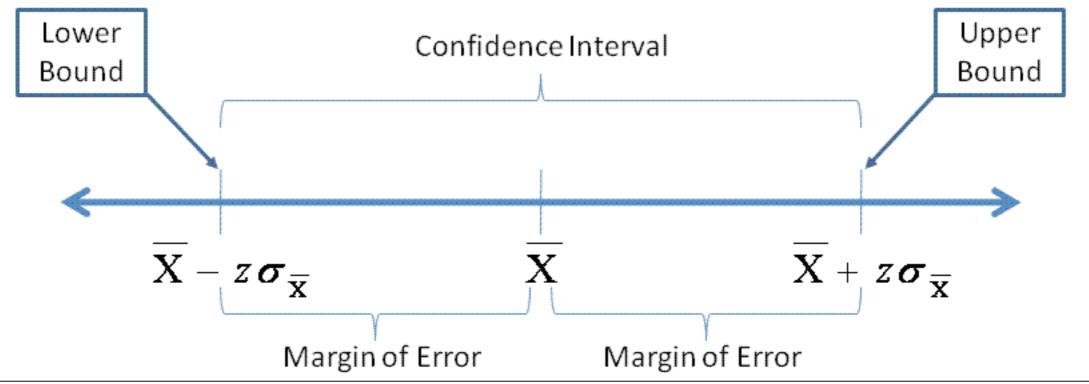
"Statistics is merely a quantization of common sense"

*"Confidence intervals consist of a range of values (interval) that act as good estimates of the unknown population parameter."* 

It is thus a way of giving a range where the true parameter value probably is.

A very simple confidence interval for a Gaussian distribution can be constructed as: (z denotes the number of sigmas wanted)





Confidence intervals are constructed with a certain **confidence level C**, which is roughly speaking the fraction of times (for many experiments) to have the true parameter fall inside the interval:

$$Prob(x_{-} \le x \le x_{+}) = \int_{x_{-}}^{x_{+}} P(x)dx = C$$

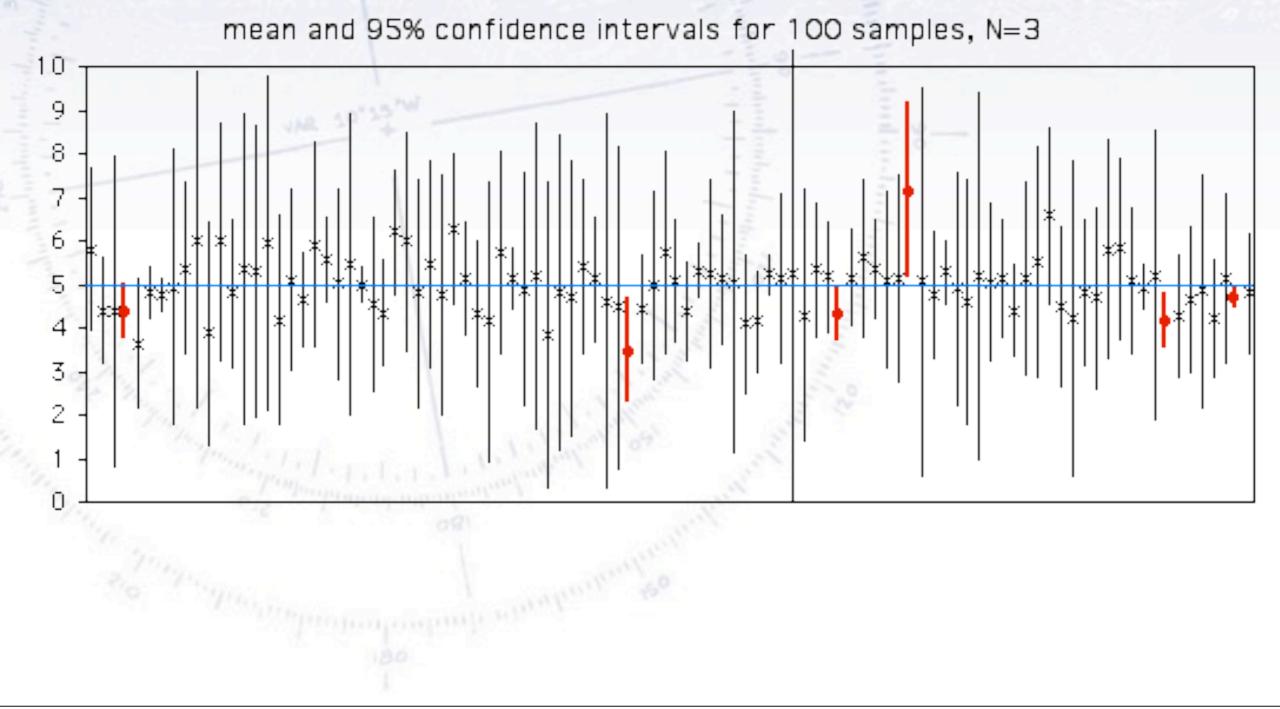
Typically, C = 95% (thus almost  $2\sigma$ ), but 90% and 99% are also used occasionally.

#### There is a choice as follows:

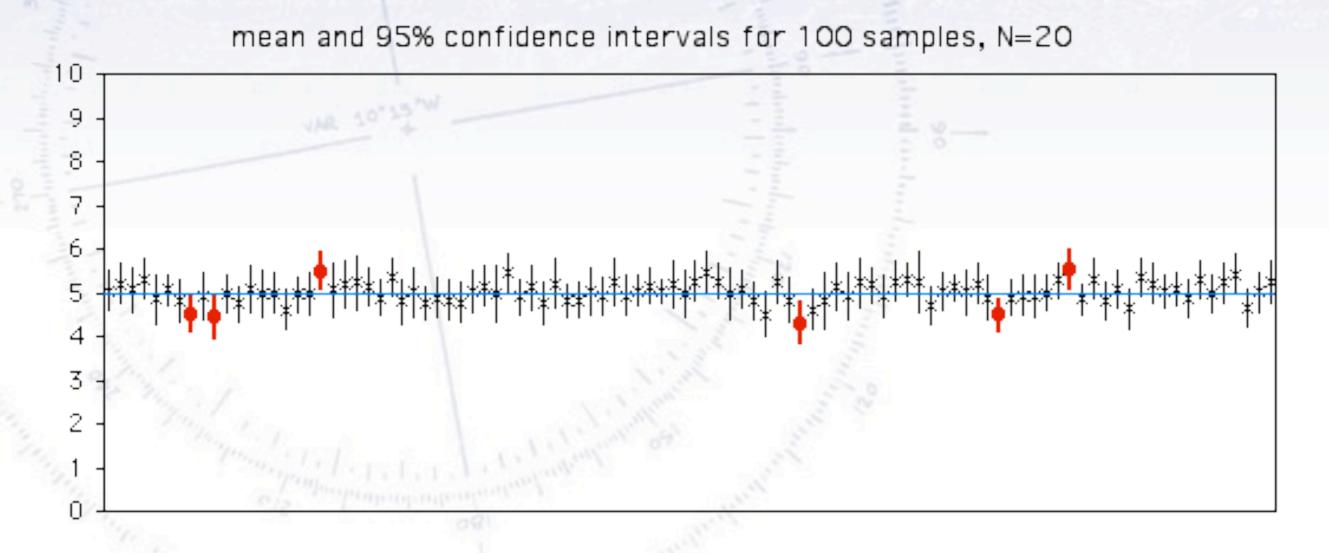
- 1. Require symmetric interval (x+ and x- are equidistant from  $\mu$ ).
- 2. Require the shortest interval (x + -x is a minimum).
- 3. Require a central interval (integral from x- to  $\mu$  is the same as from  $\mu$  to x+).

For the Gaussian, the three are equivalent! Otherwise, 3) is usually used.

The confidence interval does not ALWAYS include the true value - only C fraction.



The confidence interval does not ALWAYS include the true value - only C fraction.



...and higher statistics does not help you!

# Example

from cosmology

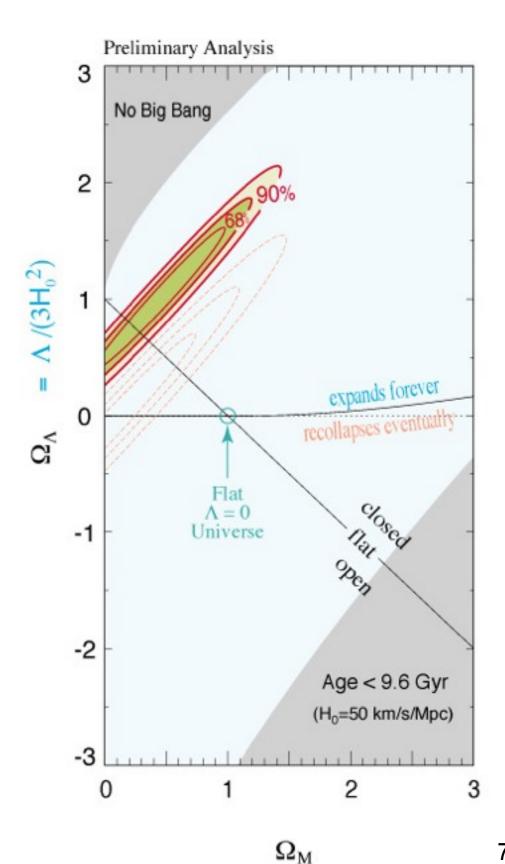
Perlmutter, et al., Nature (1998) 3 No Big Bang 2 Superit 5 supernovae A /(3H<sub>0</sub><sup>2</sup>) 00% 0  $\Omega_{\Lambda}$ -2 Age < 9.6 Gyr (H<sub>o</sub>=50 km/s/Mpc) -3 0 2 3

 $\Omega_{\rm M}$ 

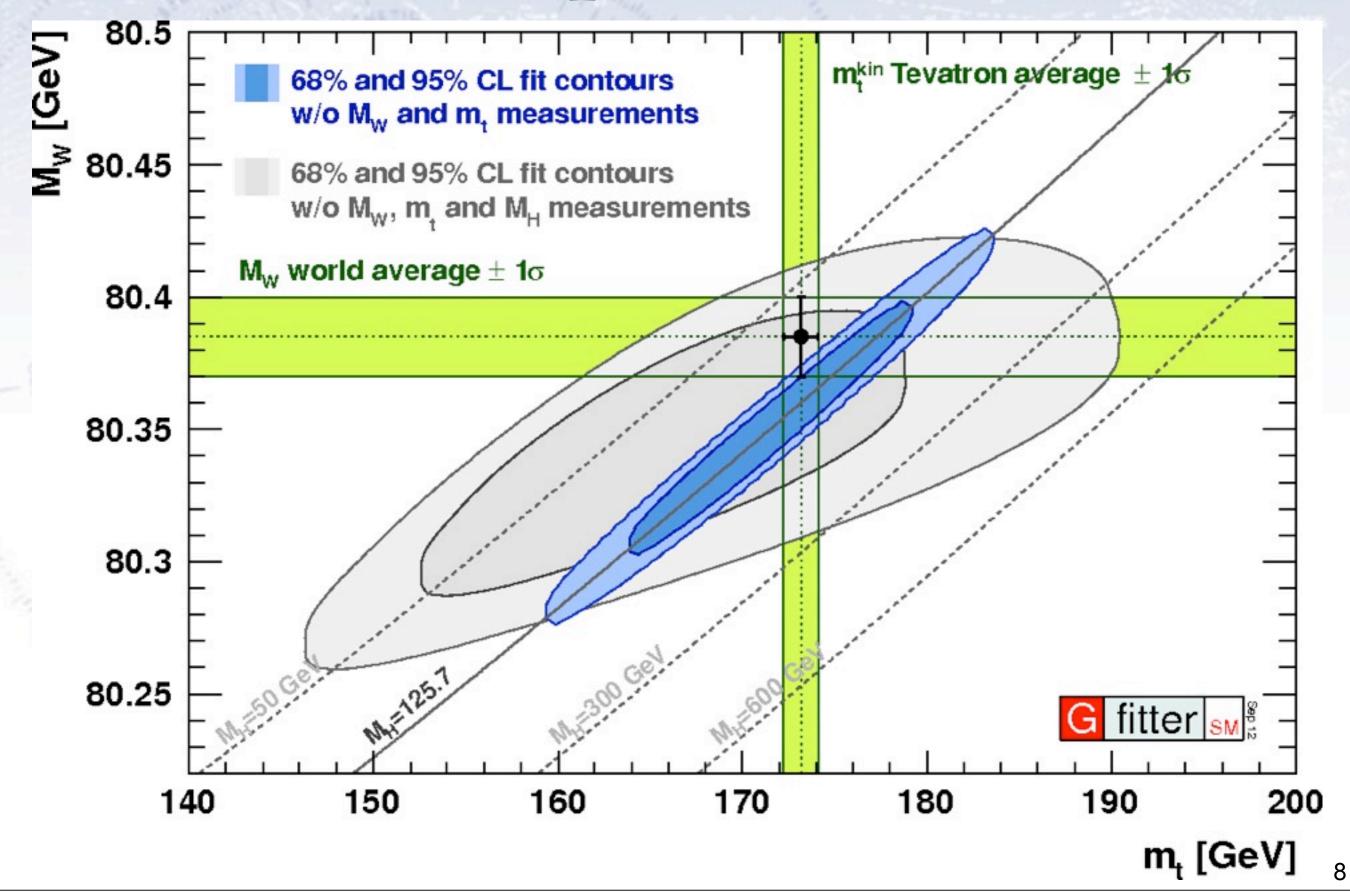
Results:  $\Omega$  vs  $\Lambda$ 

from 6 supernovae

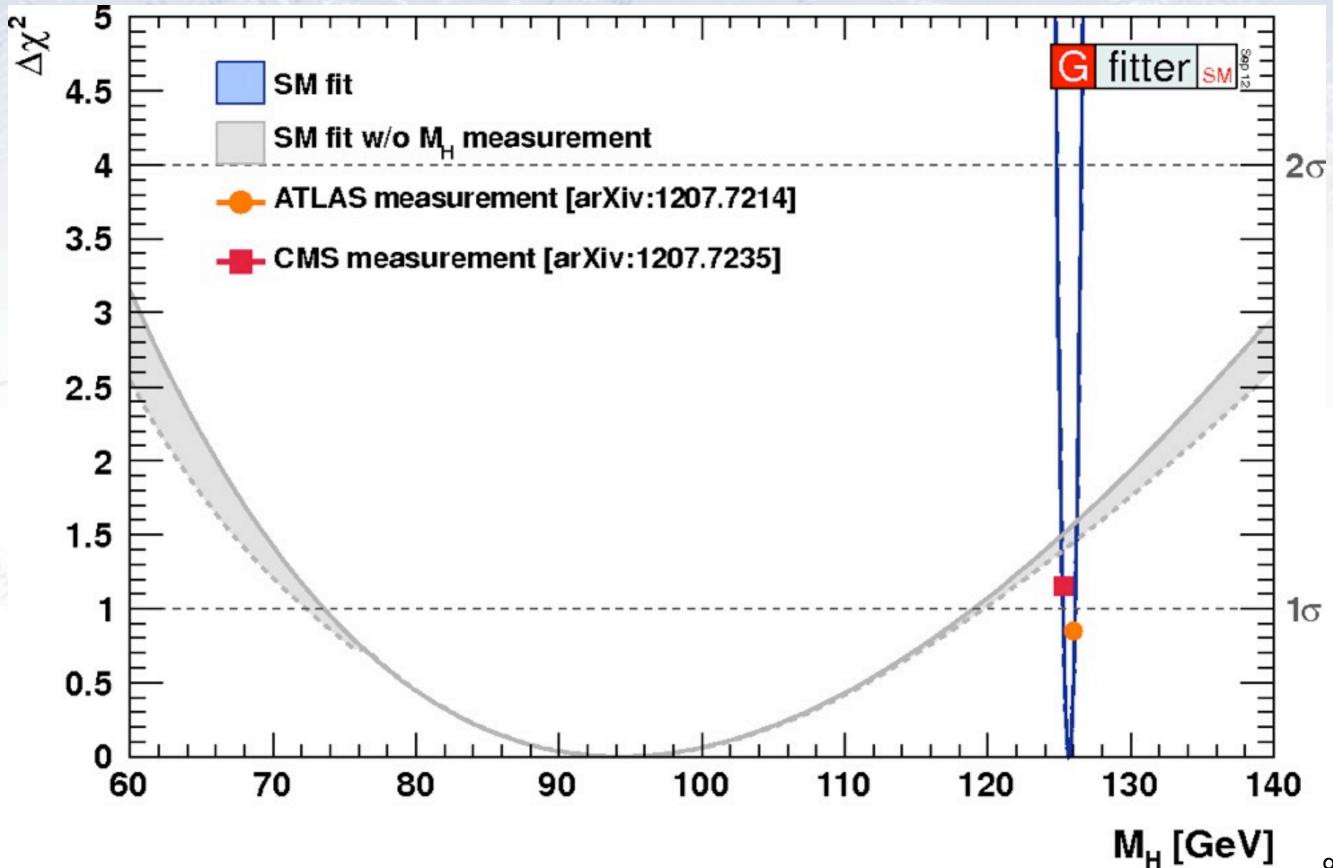
#### Results: $\Omega$ vs $\Lambda$ from 40 supernovae



#### Example from particle physics



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# **Confidence limits**

### **Confidence limits**

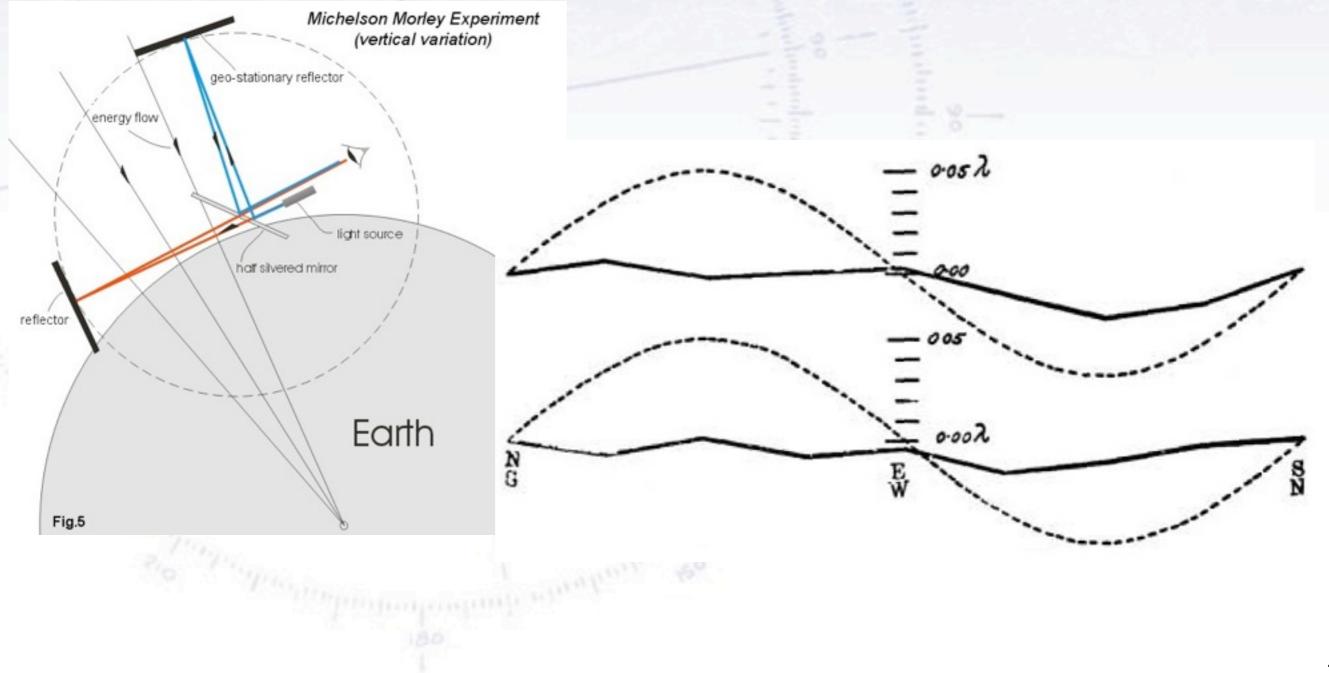
Imagine that you do an experiment to search for an unknown but predicted phenomenon (aether, planet Vulcan, dark energy, Higgs particle, etc.), and that find **nothing**!

Reporting this result, you wish to state *what you would have discovered, if it had been there,* i.e. something along the lines:

"If the aether had affected the speed of light by X%, we would have seen the effect with 95% confidence".

This is a **confidence limit** (much like a one-sided confidence interval).

In the case of Michelson-Morley, a limit could be set on the "degree of dragging" of the aether (though they didn't do this, as statistics was still in its infancy!).



### **Confidence limits - Poisson**

Poisson statistics is a neat special case, perhaps best explained by numbers:

nobs	lower limit a			upper limit b		
	$\alpha = 0.1$	$\alpha = 0.05$	$\alpha = 0.01$	$\beta = 0.1$	$\beta = 0.05$	$\beta = 0.01$
0	-	-		2.30	3.00	4.61
1	0.105	0.051	0.010	3.89	4.74	6.64
2	0.532	0.355	0.149	5.32	6.30	8.41
3	1.10	0.818	0.436	6.68	7.75	10.04
4	1.74	1.37	0.823	7.99	9.15	11.60
5	2.43	1.97	1.28	9.27	10.51	13.11
6	3.15	2.61	1.79	10.53	11.84	14.57
7	3.89	3.29	2.33	11.77	13.15	16.00
8	4.66	3.98	2.91	12.99	14.43	17.40
9	5.43	4.70	3.51	14.21	15.71	18.78
10	6.22	5.43	4.13	15.41	16.96	20.14

Table 9.3 Poisson lower and upper limits for nabe observed events.

<u>Example:</u> If you in a day observe 0 red cars on Blegdamsvej, you can with 95% confidence say that there are fewer than 3.00 pr. day.